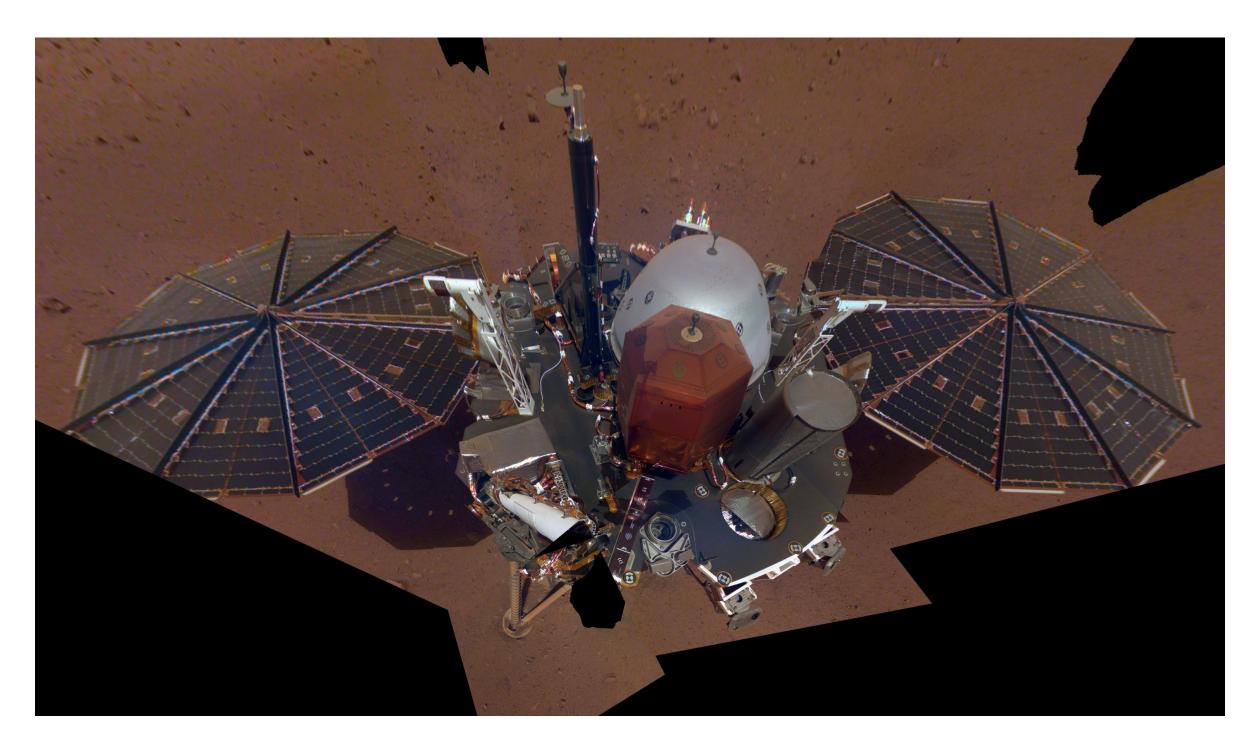
Spin state and deep interior structure of Mars from InSight radio tracking

Attilio Rivoldini Observatoire Royal de Belgique

Sébastien Le Maistre^{1,2*}, Alfonso Caldiero^{1,2}, Marie Yseboodt¹, Rose-Marie Baland¹, Mikael Beuthe¹, Tim Van Hoolst^{1,3}, Véronique Dehant^{1,2}, William Folkner⁴, Dustin Buccino⁴, Daniel Kahan⁴, Jean-Charles Marty⁵, Daniele Antonangeli⁶, James Badro⁷, Mélanie Drilleau⁸, Alex Konopliv⁴, Marie-Julie Péters¹, Ana-Catalina Plesa⁹, Henri Samuel⁷, Nicola Tosi⁹, Mark Wieczorek¹⁰, Philippe Lognonné⁷, Mark Panning⁴, Suzanne Smrekar⁴, W. Bruce Banerdt⁴.

¹Royal Observatory of Belgium, Brussels, Belgium;²UCLouvain, Louvain-la-Neuve, Belgium; ³Institute of Astronomy, KU Leuven, Leuven, Belgium; ⁴Jet Propulsion Laboratory, California Institute of Technology, Pasadena, USA.; ⁶Centre national d'Études Spatiales, Toulouse, France. ⁶ IMPMC, Sorbonne Université, MNHN, CNRS, Paris, France. ⁷ Université de Paris, Institut de Physique du Globe de Paris, CNRS, Paris, France; ⁸ Institut Supérieur de l'Aéronautique et de l'Espace SUPAERO, Toulouse, France; ⁹ DLR Institute of Planetary Research, Berlin, Germany; ¹⁰ Université Côte d'Azur, Observatoire de la Côte d'Azur, CNRS, Laboratoire Lagrange, Nice, France.

The Rotation and Interior Structure Experiment



- **RISE** is together with **SEIS** and **HP3** one of the main instruments of the InSight mission
- **RISE**: determine the rotation of Mars
 - precession
 - measure the nutation of the spin axes to detect and quantify the effect of the liquid core
 - budget

• measure the rotation rate of Mars on a seasonal timescales to constrain the atmospheric angular momentum

RISE setup



- placed on the InSight platform
- Earth: Deep Space Network tracking stations (USA, Australia, Spain)
- the InSight lander and coherently retransmitted back to the DSN

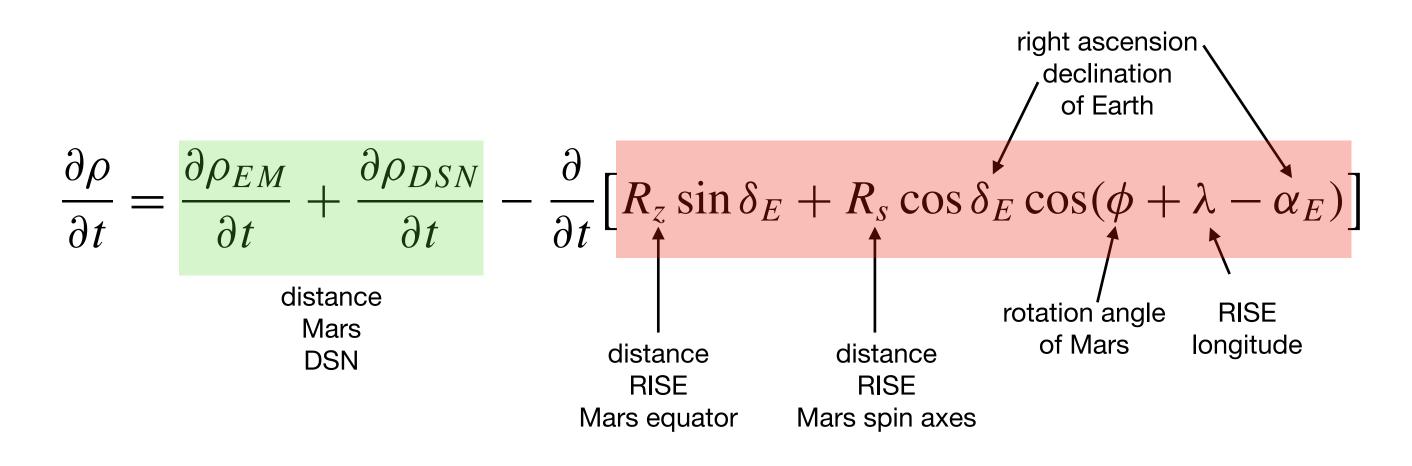


Mars: 2 medium gain antennas (east and west) and a X-band radio transponder

mesure 2-way Doppler shifts at DSN stations of a radio signal sent from the DSN to

Doppler shifts and lander positioning in space

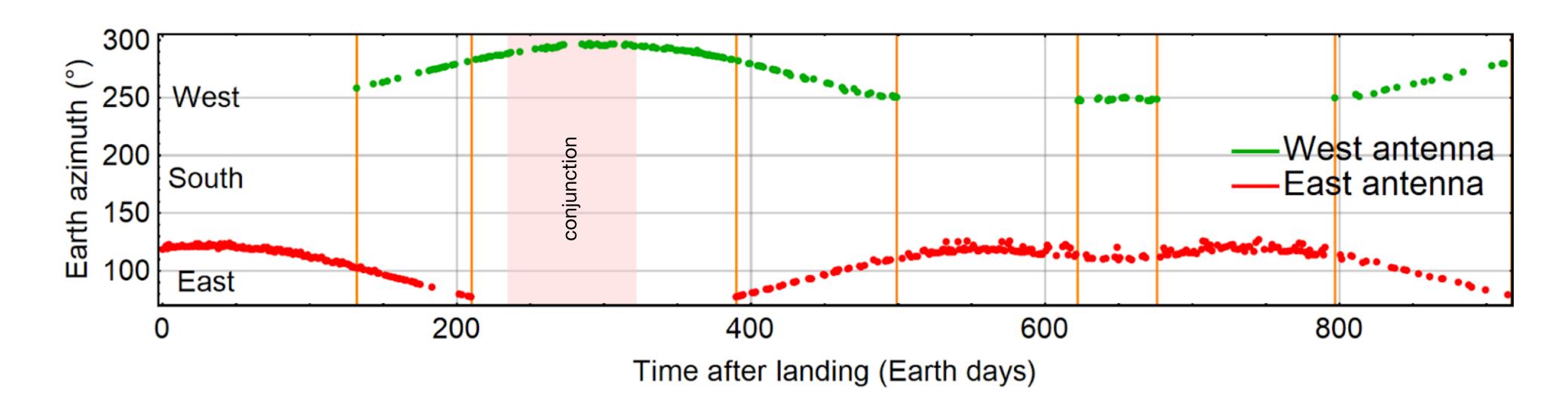
 the measured Doppler shift is proportional to the rate of change of distance between the DSN station and the lander ρ (approx.)



- know: DSN-Mars distance
- infer: lander position, rotation angle, declination, right ascension ascension

\Rightarrow precession rate and nutation determined from declination and right

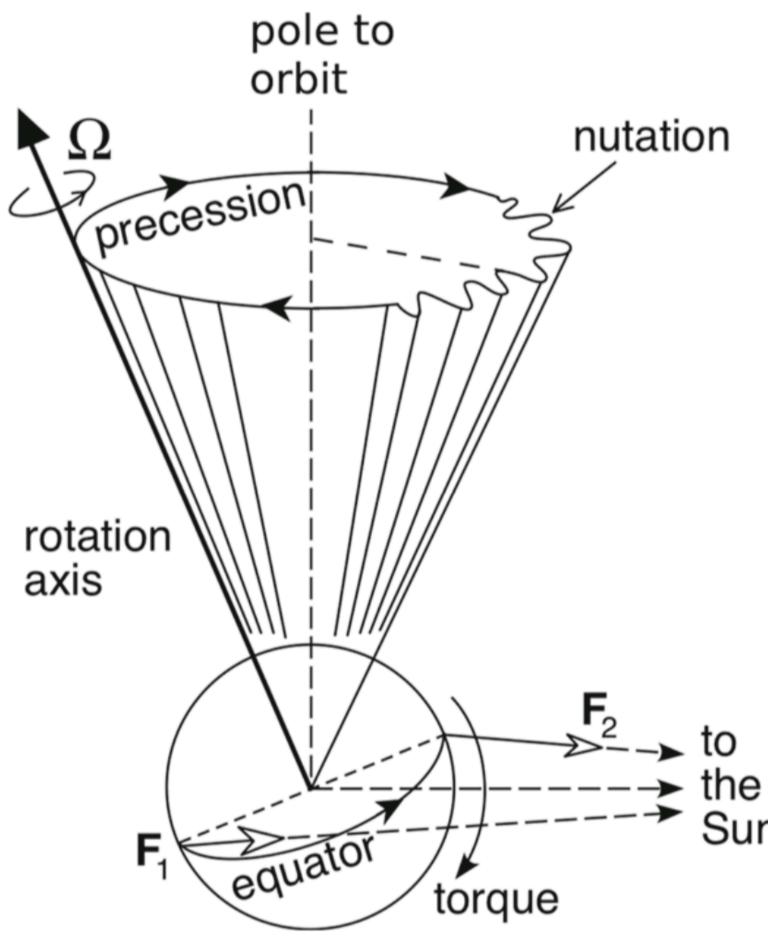
Tracking



- Earth seen in the morning with the eastward antenna and with the westward in the the afternoon (elevation 10° to 50°)
- first Martian year days: 5-7 trackings each week (60min) later reduced to 1-3 (30min)
- collected 25 0000 Doppler data points over 915 days

- the gravitational torque exerted by the Sun on the flattened rotating Mars causes a precession of the rotation axis in space (~171000 years) \Rightarrow precession rate ~ to the polar moment of inertia
- torque variations due to the relative positions between the Sun and Mars lead to periodic motions of the rotation axis, the nutations (1/ (1,2,3,4..) year) \Rightarrow lander position changes by about 10m on the surface

Precession and nutation



Nutation: interior structure

- if a planet were rigid then nutation amplitudes can be predicted very precisely from its moment of inertia and from the tidal potential (well known forcing periods)
- nutation amplitudes depend on the interior structure of Mars and in particular on the liquid core
- the relative rotation between the fluid core and solid mantle is characterised by a rotational normal mode, the Free Core Nutation
- if the FCN frequency ω_{FCN} is close to forcing frequency ω the nutation amplitude can be resonantly amplified

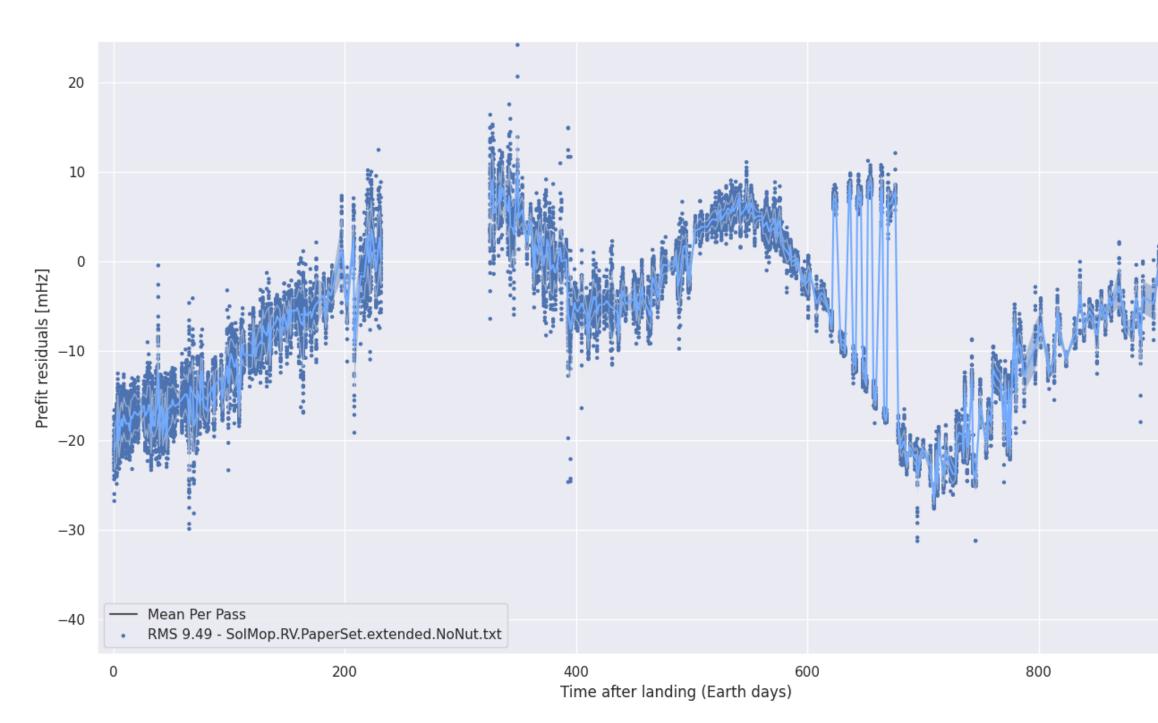
• the amplification strength $F = \frac{A_f}{A - A_f} \left(1 - \frac{A_f}{A - A_f}\right)$

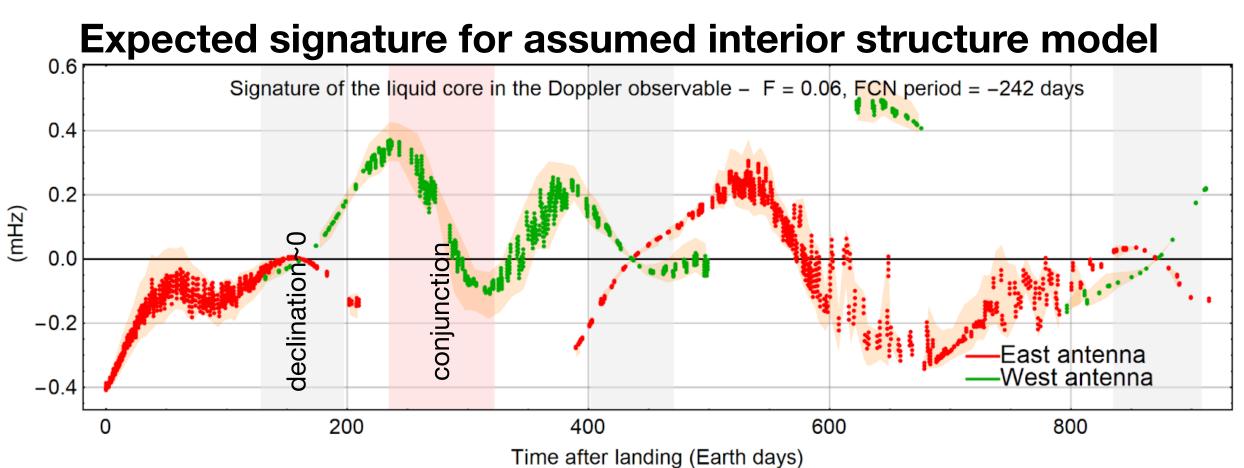
the interior structure of the planet compliances due to tidal forcing (γ) and rotation rate variation (β)

$$\left(\frac{\gamma}{e}\right)$$
 and $\omega_{FCN} = -\Omega \frac{A}{A - A_f}(e_f - \beta)$ are related to

 \Rightarrow moments of inertia of the planet (A) and core (A_f), planet (e) and core dynamic shape (e_f), core

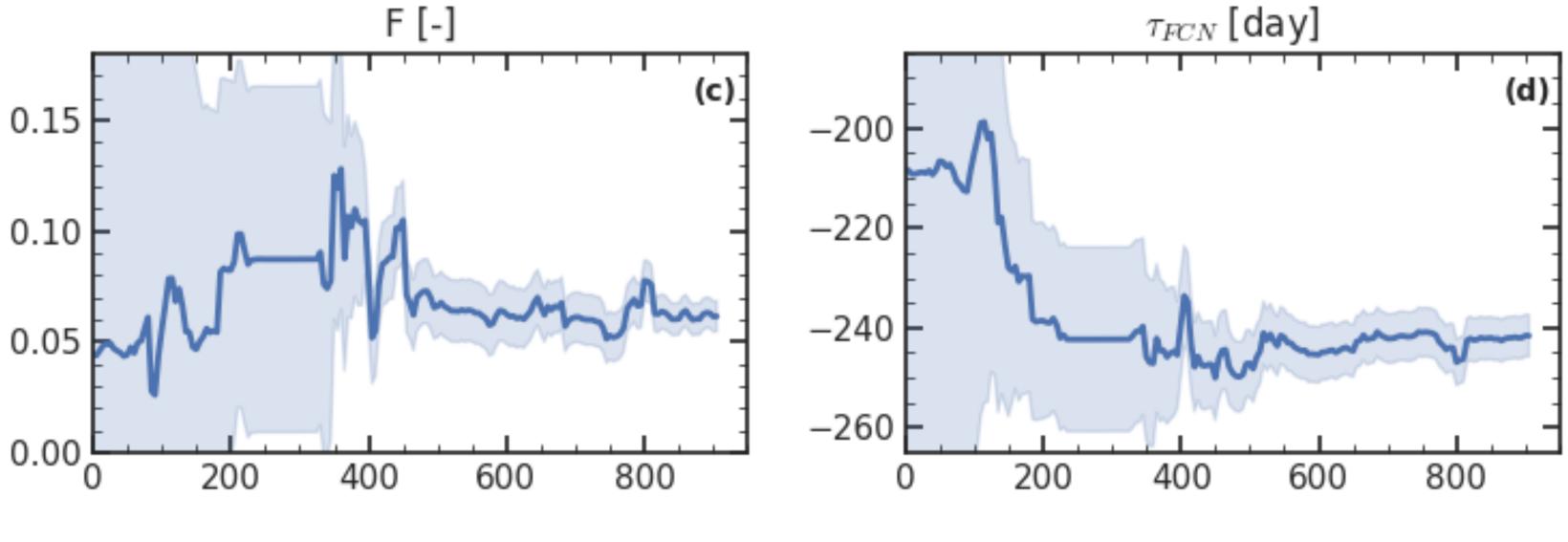
Liquid core signature and real data





- the measured doppler shift is about 20-30 mHz
- the signature of the liquid core is 2 orders of magnitude smaller
- since its periods are well known and because of data accumulation it can be determined





Time after landing (Earth days)

- amplification factor F and FCN period $\tau_{\rm FCN}$
- $F = 0.061 \pm 0.0064$ and $\tau_{\text{FCN}} = -242.25 \pm 2.7$ days
- F in expected range but $\tau_{\rm FCN}$ somewhat lower than expected

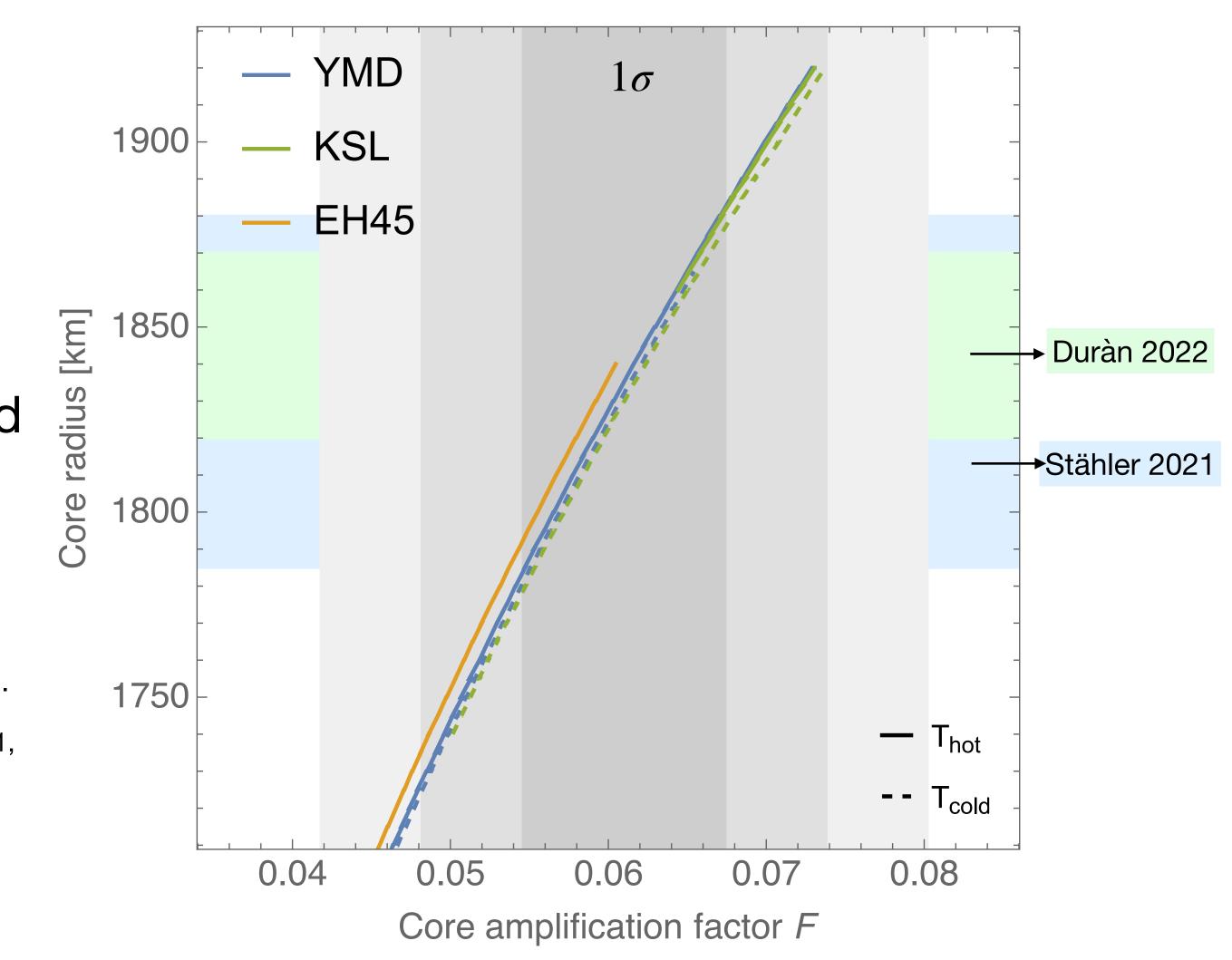
RISE results

>600 days of data are required to obtain robust estimates for the core

Interpretation: core radius

- assume 3 mantle compositions (E45 Sanloup 1999, YMD Yoshizaky 2020, KSL Khan 2022) mantle temperature end-members, average crust structure in agreement with seismic data (Knapmeyer-Endrun 2021)
- $F(\sim A_f)$ is robust with respect to T and mantle composition
- inferred core range in excellent agreement with tidal measurements (e.g. Rivoldini et al. 2011) and seismic data (Stähler et al. 2021, Duràn et al 2022)

 $\Rightarrow r_{\rm cmb} = 1825 \pm 55 \text{ km}$

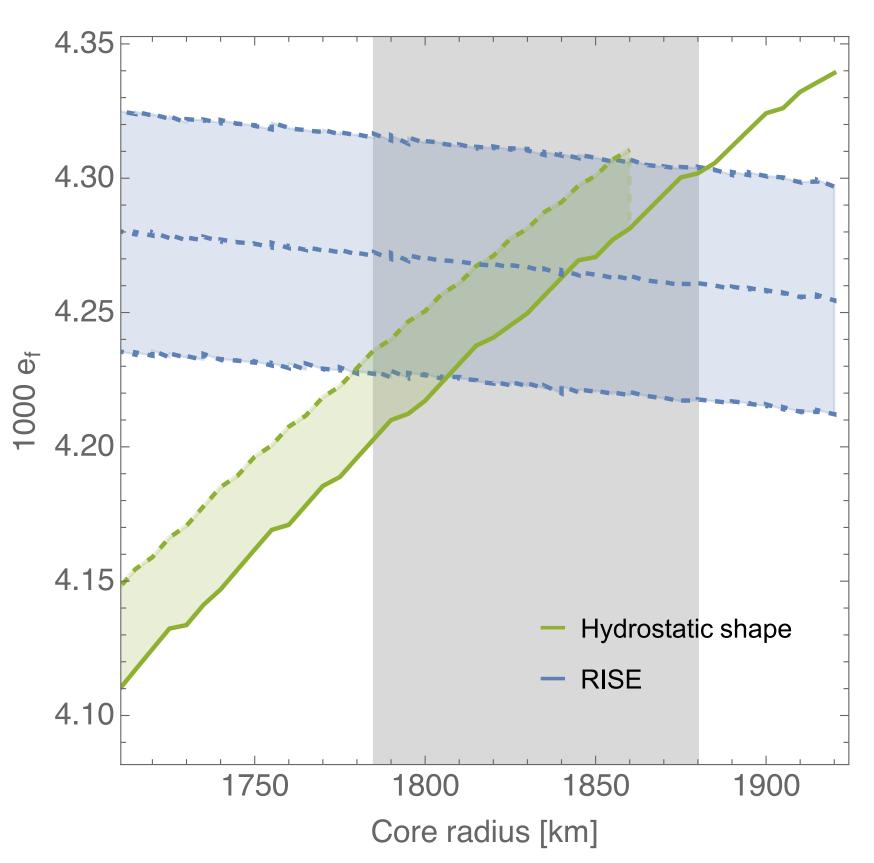


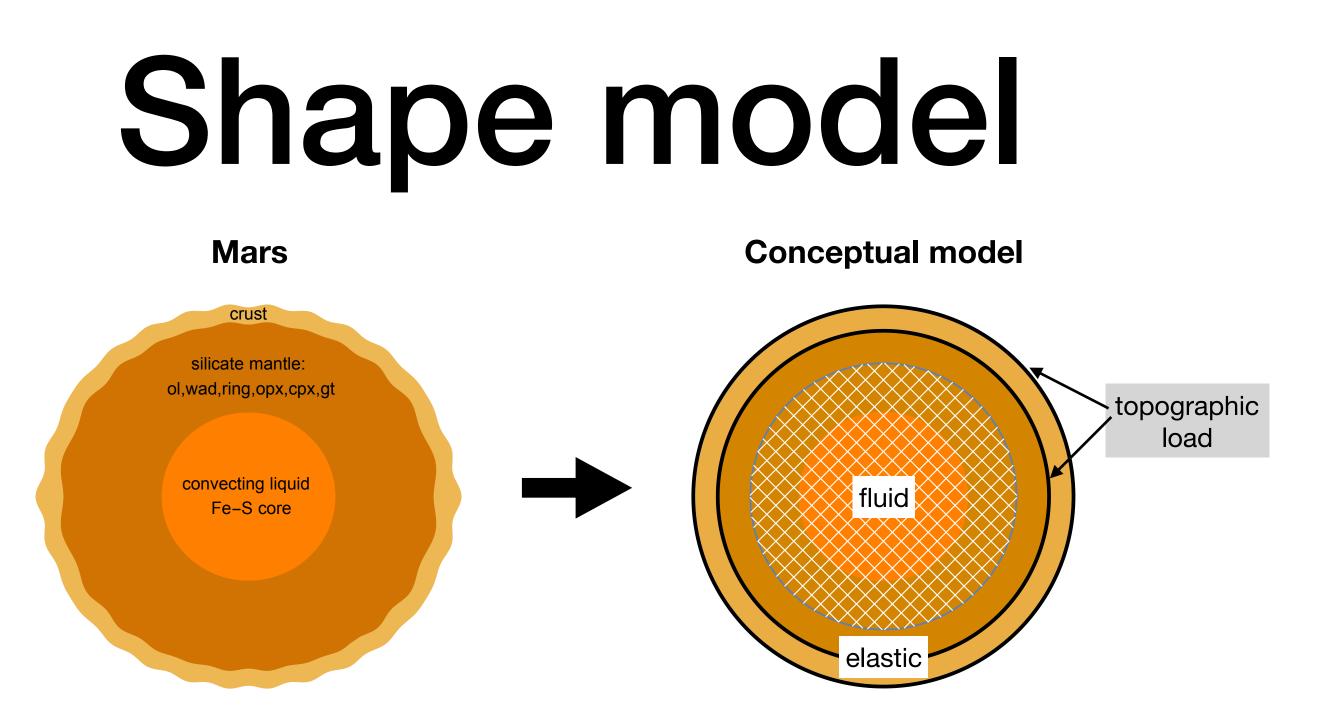
Interpretation: FCN period

• FCN frequency proportional to dynamic core shape $e_f = \frac{C_f - A_f}{A_f}$ which is directly related to the density jump at the

core mantle boundary

- RISE data implies an almost hydrostatic core shape
- **but** Mars' geometric and dynamic shape $(e = \frac{C A}{A})$ are **not hydrostatic**





- geometric and dynamic shape (deg 2) of Mars results from rotation, mass **boundary** (see also Zharkov 2009, Wieczorek 2019)

anomaly induced by the surface topography, and internal mass anomalies placed deep within the planet (Moho, bottom of the lithosphere, core-mantle

internal loads are specified to match geometric and dynamic shape of Mars

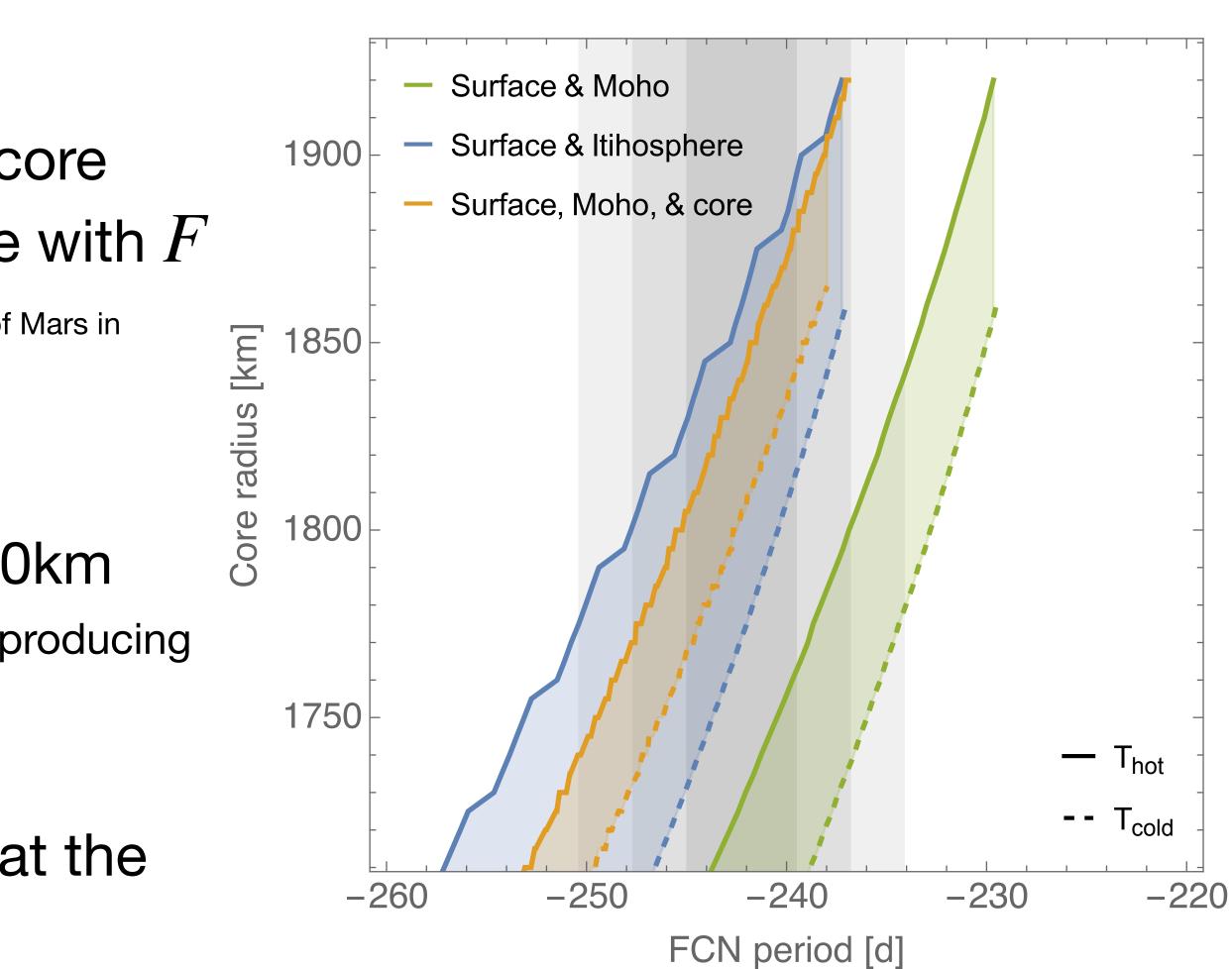
Interpretation: FCN period

Models agree with the FCN period if:

- a load is placed at the Moho, but the core radius of those models does not agree with ${\cal F}$

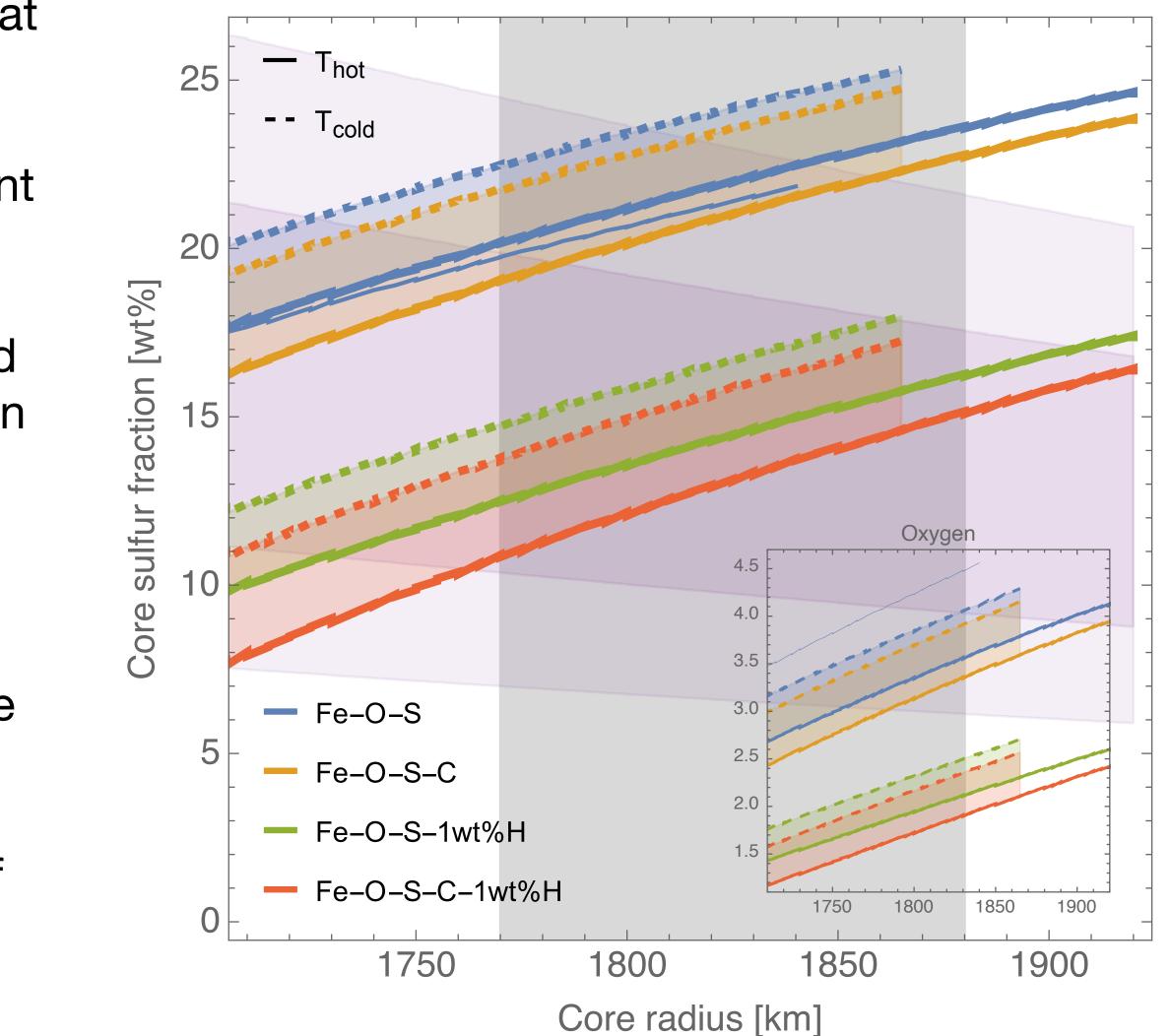
(an equivalent approach is used to deduce the average crust thickness of Mars in Knapmeyer-Endrun 2021)

- a load is placed at the bottom of the lithosphere, but the lithosphere is >550km
 ⇒ requires that the crust is highly enriched in heat producing elements compared to the mantle
- a load placed at shallower depth and at the core-mantle boundary



Interpretation: core composition

- candidate light elements that are siderophile at core forming conditions: O, S, C, H
- amounts of light elements are not independent and dependent on the bulk chemistry
- use core formation model to relate O to S and mantle composition and set C to its saturation
 - O increases with S and mantle FeO
 - C decreases with increasing S (C≤2wt%)
- models without H are only possible if the core radius is at the lower end of our estimate
- RISE compatible models require >0.6wt% of H if S is in agreement with geochemical constraints



Conclusions

- core
- RISE data constrain the moment of inertia of the core, the density jump at the core mantle boundary, and the shape of the core
- 1.5±0.3wt% C if 1wt% H is assumed in the core
- chemical anomalies

• the measured nutation and the detection of the FCN normal mode confirm the liquid state of the

• the core radius is in excellent agreement with estimates obtained from tides and seismic data

• RISE data and geochemical constraints require a core with 2.1±0.3wt% O, 14.5±1.5wt% S, and

• the measured FCN period requires an almost hydrostatic core shape, such a core shape can be explained by deep seated mass anomalies within the mantle that originate form thermal or

• one supplemental Earth year of RISE data will decrease the uncertainties on F and $au_{
m FCN}$ by 5%