

Solar-Terrestrial
Center of
Excellence



Royal Observatory
of Belgium

Near Real-Time Monitoring of the Solar Activity Impact on European Region from the EPN Data

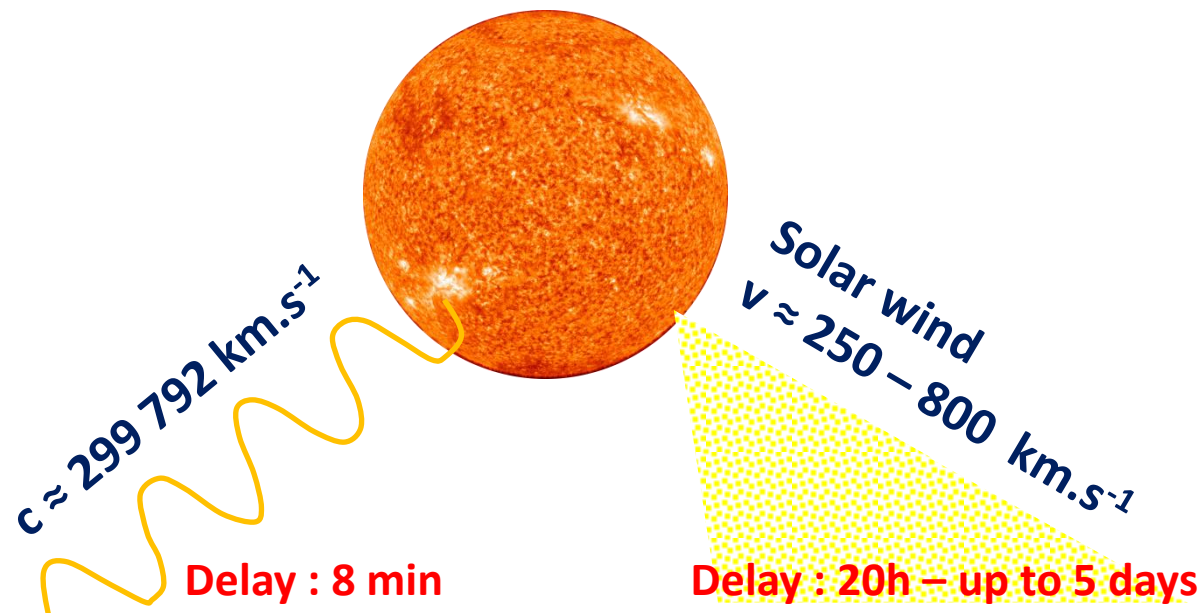
J.-M. Chevalier and N. Bergeot



Brussels, Belgium
October 25-26, 2017

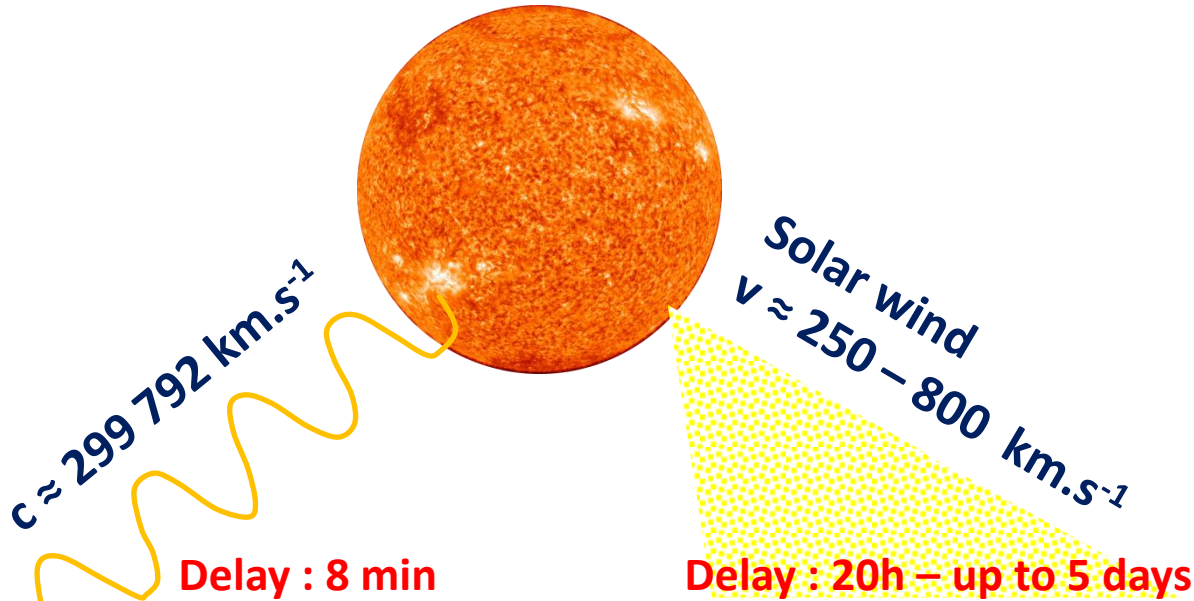


ROYAL OBSERVATORY
OF BELGIUM



Electromagnetic Radiations
EUV X-rays Radio Waves

Energetic Electrically charged Particles
Inter-planetary Magnetic Field
Protons & Electrons

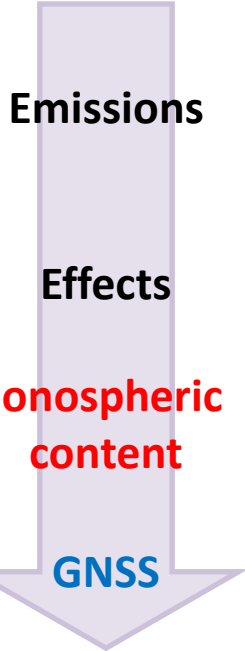


Electromagnetic Radiations
EUV X-rays Radio Waves

Photo-ionisation

Ionospheric Ne Δ

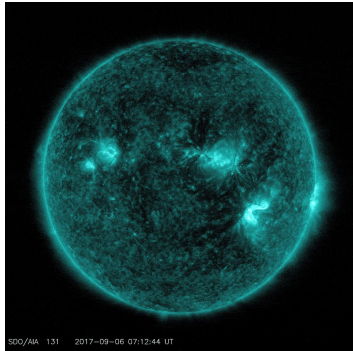
Radio signal delay



Energetic Electrically charged Particles
Inter-planetary Magnetic Field
Protons & Electrons

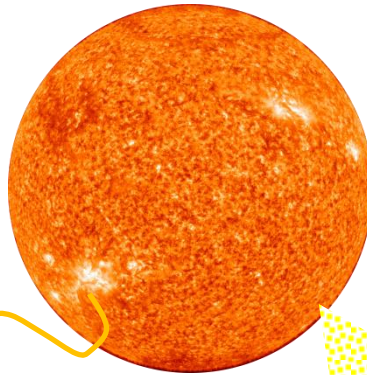
Introduction

The Sun, Source of Space Weather



NASA

Solar Flare



$c \approx 299\,792\text{ km}\cdot\text{s}^{-1}$
 Delay : 8 min

Solar wind
 $v \approx 250 - 800\text{ km}\cdot\text{s}^{-1}$
 Delay : 20h – up to 5 days

Electromagnetic Radiations
 EUV X-rays Radio Waves

Photo-ionisation

Solar Radio Burst

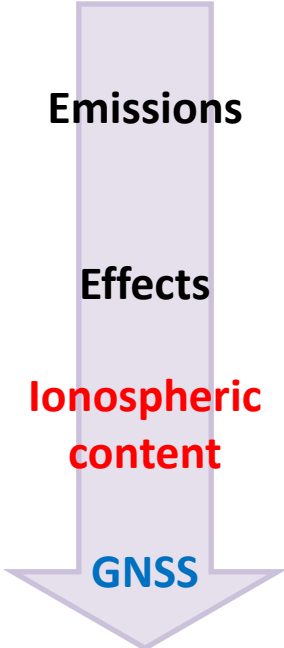
Ionospheric $N_e \Delta$ – Scintillations

Signal reception fade (\nearrow noise)

Radio signal delay-
 \searrow Signal reception

Signal reception
 fade (\nearrow noise)

Energetic Electrically charged Particles
 Inter-planetary Magnetic Field
 Protons & Electrons

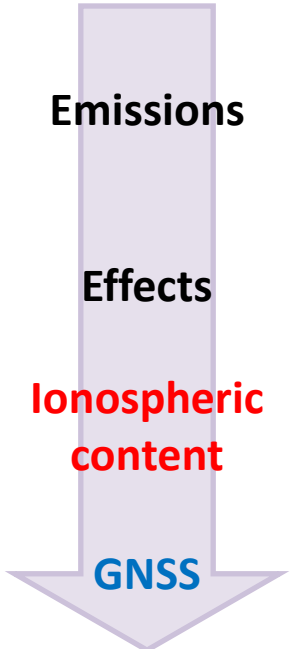


Emissions

Effects

Ionospheric content

GNSS



| Electromagnetic <u>Radiations</u> | |
|--|-------------|
| EUV X-rays | Radio Waves |

Photo-ionisation

Solar Radio Burst

Ionospheric Ne Δ – Scintillations

Radio signal delay - \searrow Signal reception

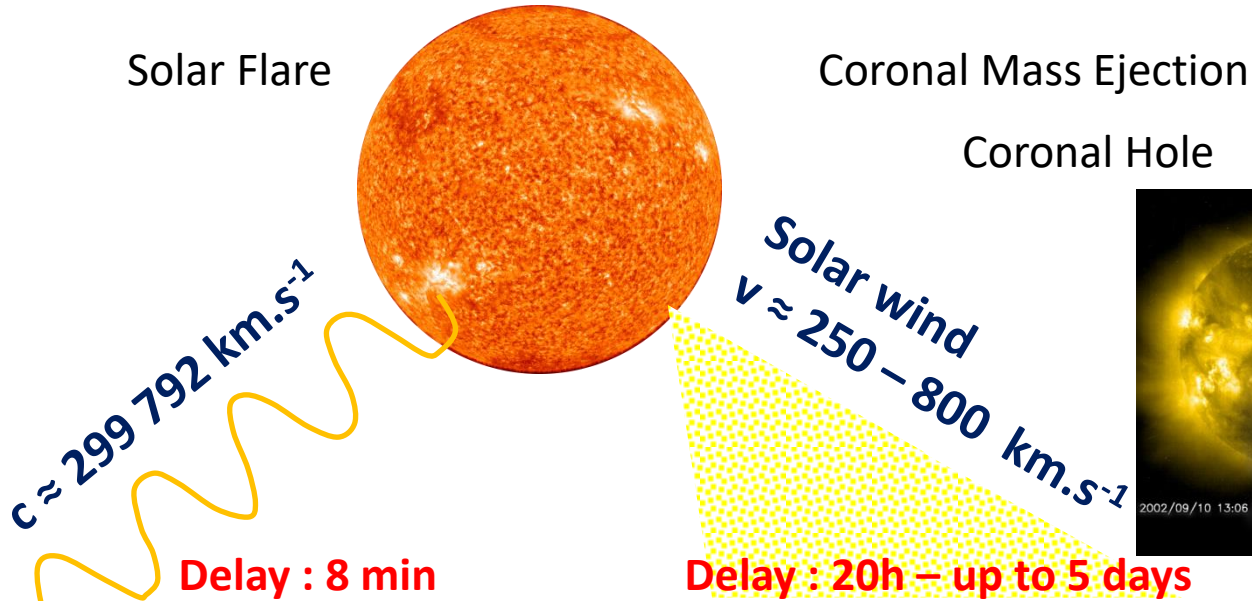
Signal reception fade (\nearrow noise)

| Energetic Electrically charged <u>Particles</u> |
|--|
| Inter-planetary Magnetic Field Protons & Electrons |

Geomagnetic Storm
Ionisation & Recombination - Aurora

Ionospheric Ne disturbances - Scintillations

Radio signal delay - \searrow Signal reception



| Electromagnetic <u>Radiations</u> | |
|--|-------------|
| EUV X-rays | Radio Waves |

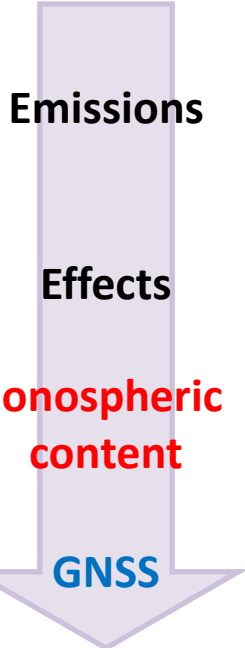
Photo-ionisation

Solar Radio Burst

Ionospheric Ne Δ – Scintillations

Radio signal delay -
↘ Signal reception

Signal reception fade (↗noise)

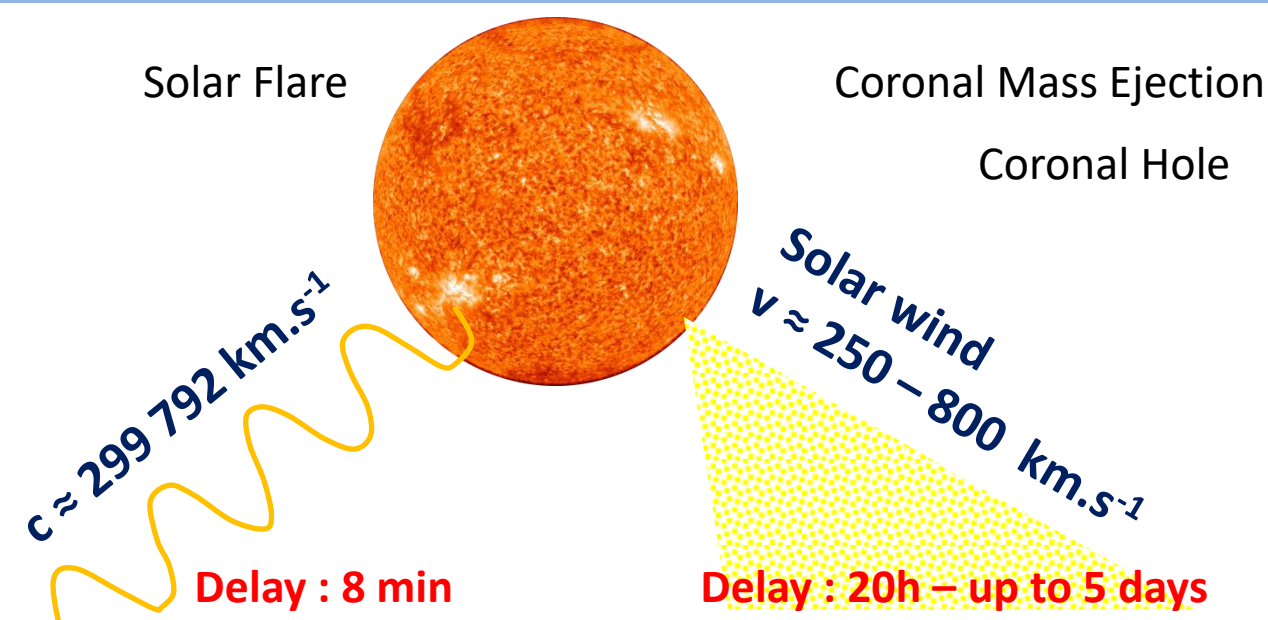


| Energetic Electrically charged <u>Particles</u> |
|--|
| Inter-planetary Magnetic Field Protons & Electrons |

Geomagnetic Storm
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Ionospheric Ne disturbances - Scintillations

Radio signal delay -
↘ Signal reception



Electromagnetic Radiations
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Photo-ionisation Solar Radio Burst

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Emissions

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GNSS

Energetic Electrically charged Particles
 Inter-planetary Magnetic Field
 Protons & Electrons

Geomagnetic Storm
 Ionisation & Recombination - Aurora

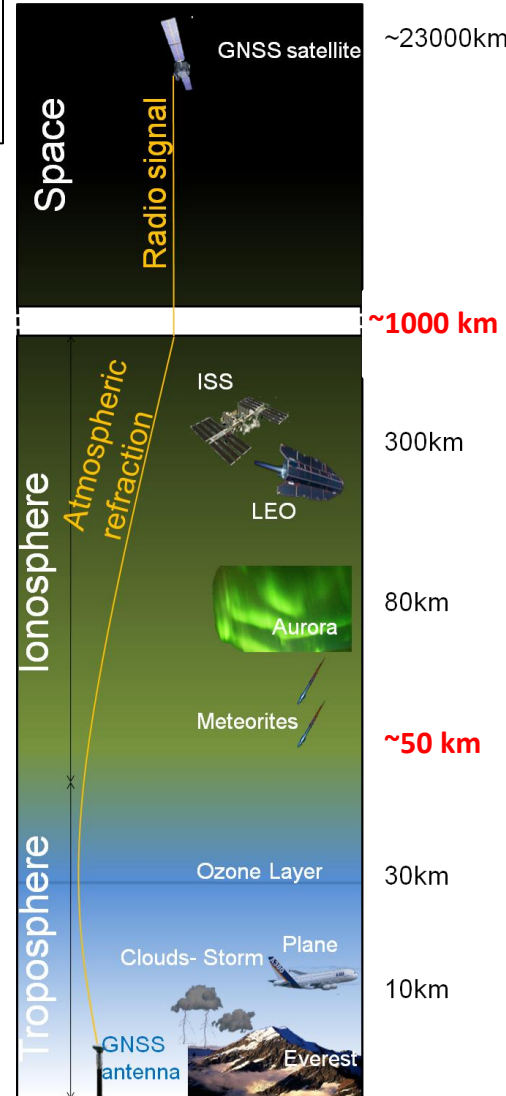
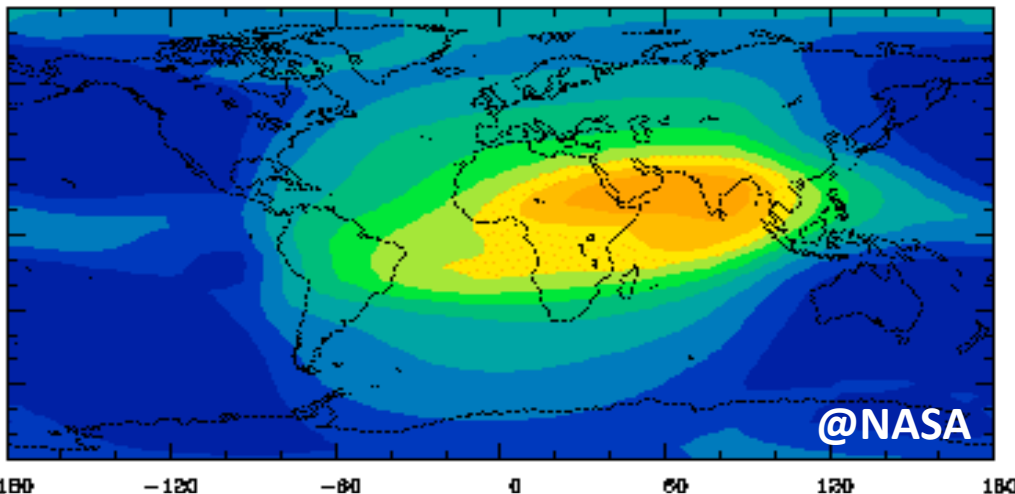
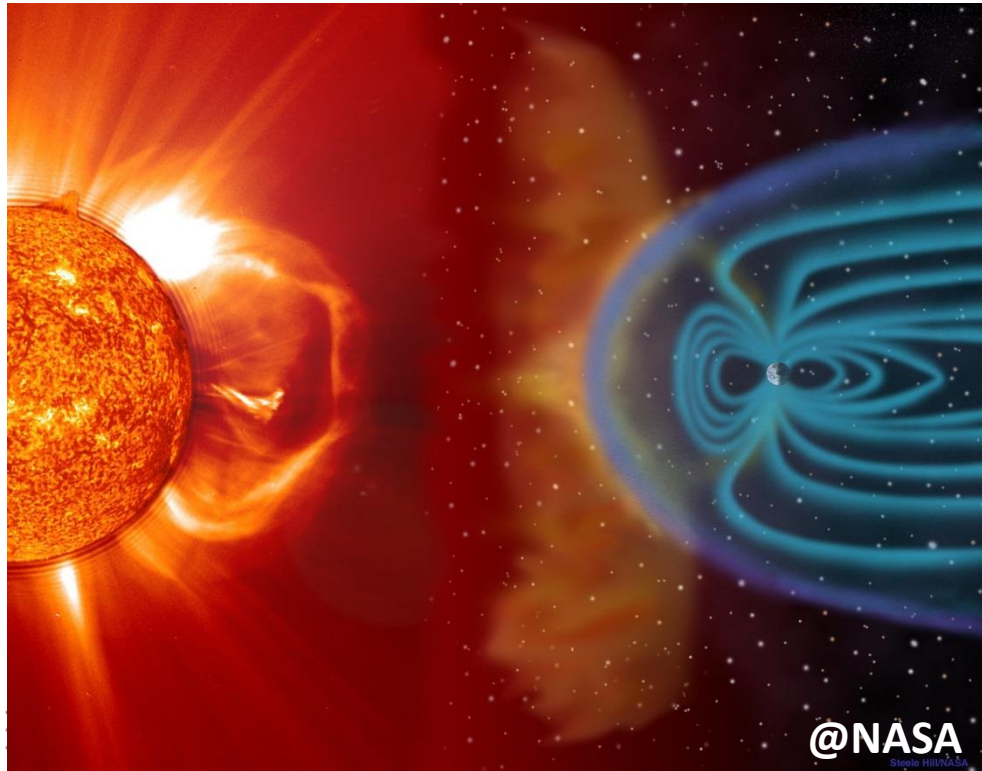
Ionospheric Ne disturbances - Scintillations

Radio signal delay -
 ↘ Signal reception

Ionised media

Plasma:

$$e^- = ion^+$$



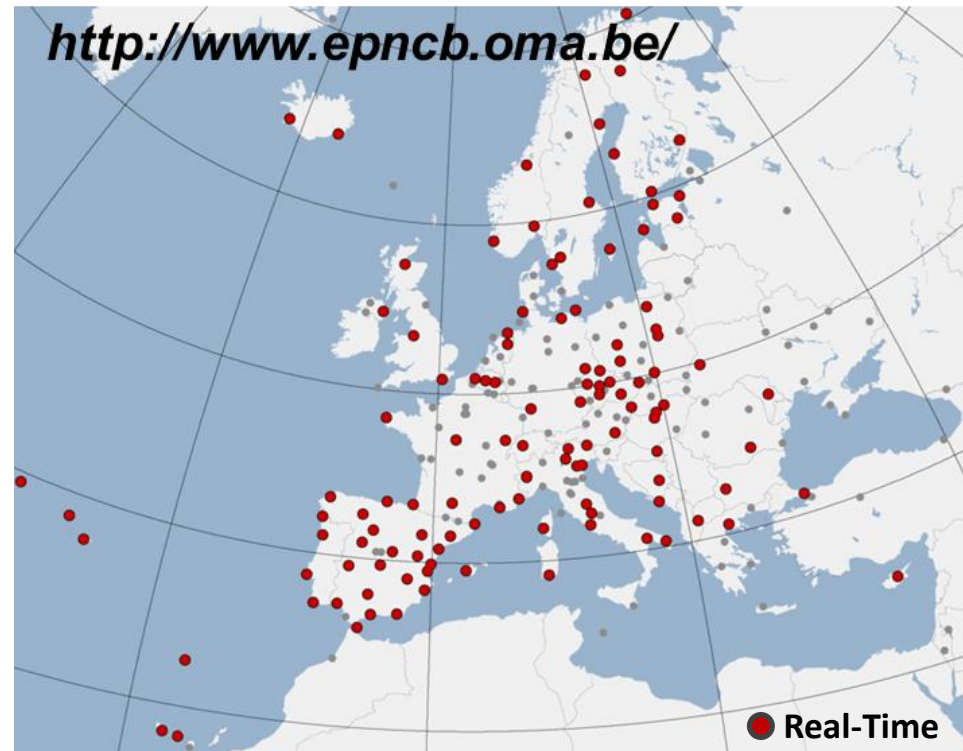
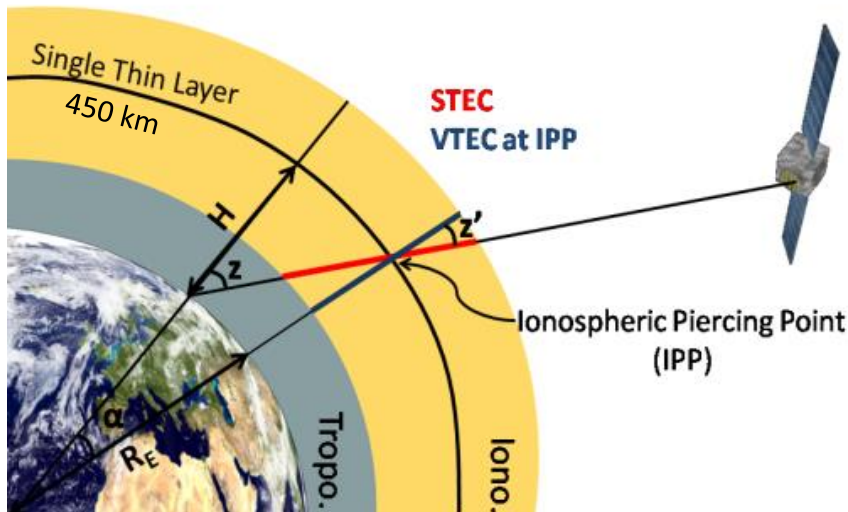
Electrically charged media affects the radio-wave propagation (depending on the frequency)

=> ionospheric delay $I_{1,2}$

Pseudorange $P_{1,2} = \rho_{1,2} - c\Delta t_{rec} + c\Delta t_{sat} + I_{1,2} + Tr + \delta_{1,2}^P + \epsilon_{1,2}^P$

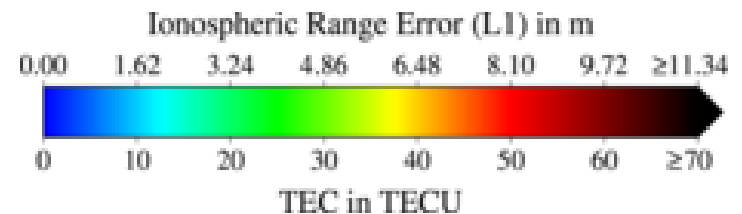
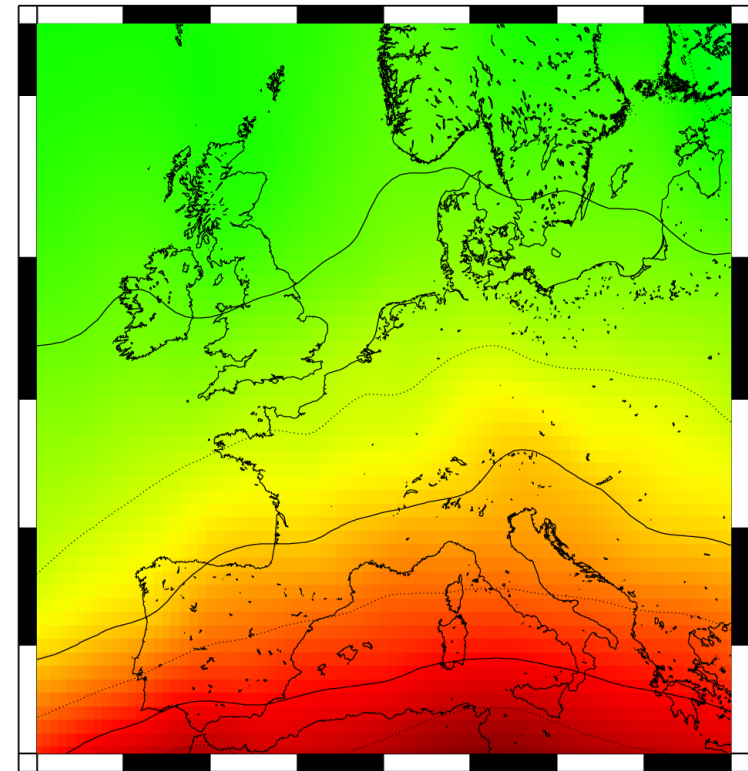
Ionospheric delay $I_{1,2} = 40.3 \frac{STEC}{f_{1,2}^2}$

Slant Total Electron Content $STEC = \int_R^S N_e d$



- Input data: real-time data from the EPN (~150 stations)
- Near real-time TEC maps over Europe + variability
 - since 2012

| | |
|-----------------|---|
| Sampling rate | : 15 min |
| Grid extent | : Long W15° / E25° : Lat N35° / N62° |
| Grid resolution | : 0.5°x0.5° |
| Latency | : ~3 minutes |



Last Ionospheric Events

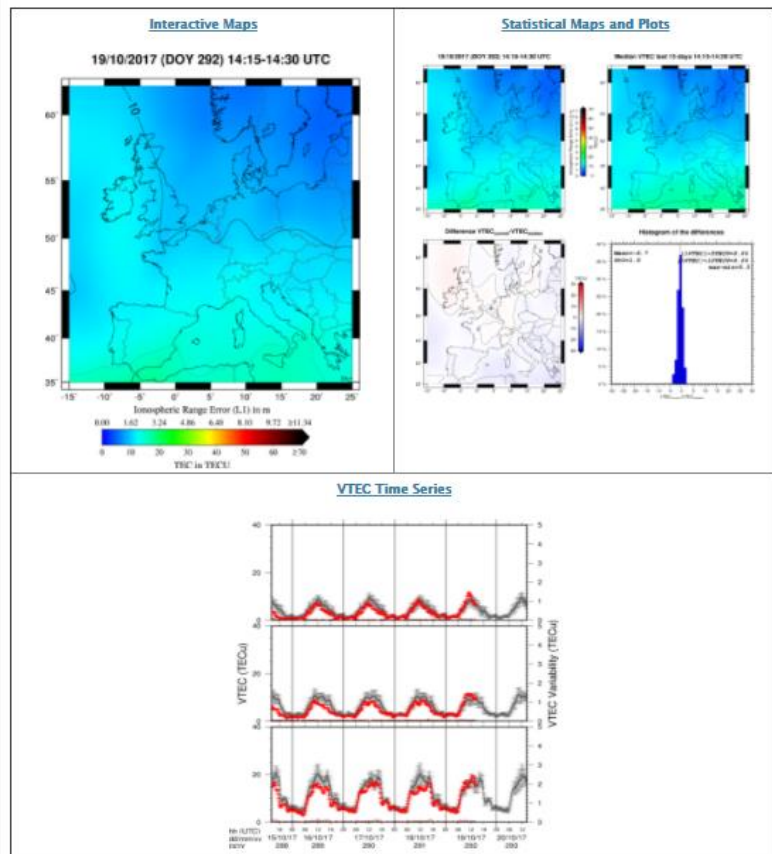
- 2017-09-07 : Ionospheric activity due to Solar Flare ([more here](#))
- 2017-09-06 : Space weather event due to Solar Radio Burst ([more here](#))
- 2017-08-21 : Ionospheric activity due to Geomagnetic Activity ([more here](#))
- ... [more events here](#)

Near-Real Time Ionospheric Products

Vertical Total Electron Content (VTEC) estimated in Near Real-Time (NRT) every 15 minutes from EUREF Permanent Network (EPN) GPS data.

[More...](#)

- [Interactive Maps](#): display animated VTEC maps (movie) for a requested period and VTEC value at a given location and time. (4-5 sec to load).
- [Statistical Maps and Plots](#): statistics to compare the ionosphere for a requested time with respect to the 15 previous days.
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- [Data](#) are publicly available in IONEX format at <http://gnss.oma.be/gnss/products/IONEX/>. We request that users include a citation or an acknowledgment when using ROB VTEC data or products results in a publication. See [disclaimer and copyright](#) for more information.



Solar Radio Burst Warnings for GNSS Applications in Europe

Solar Radio Bursts (SRB) emitted at the GNSS frequencies can affect the GNSS signal reception. To detect such event, a [near-real time SRB warning system](#) with a 4-level index was set in Europe using the real-time EUREF Permanent Network.

Monitoring based on Real-Time EPN Data

www.gnss.be

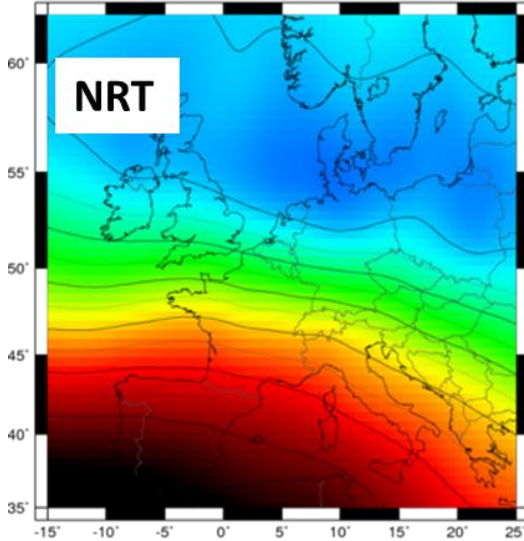
1) Ionospheric Total Electron Content (TEC)

- Interactive TEC maps
- Statistical TEC maps
- TEC Time Series at 3 locations (North-Brussels-South)
- TEC Data (IONEX) [ftp://gnss.oma.be/gnss/products/IONEX/](http://gnss.oma.be/gnss/products/IONEX/)
- Event Description

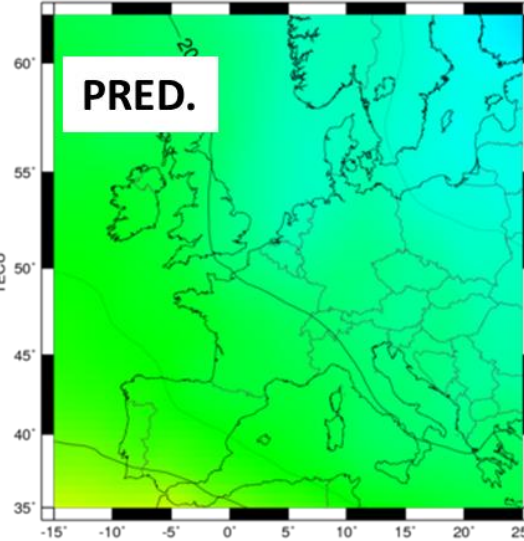
2) Solar Radio Burst (SRB)

- Warning System
- Event Description

17/03/2015 (DOY 076) 18:15-18:30 UTC

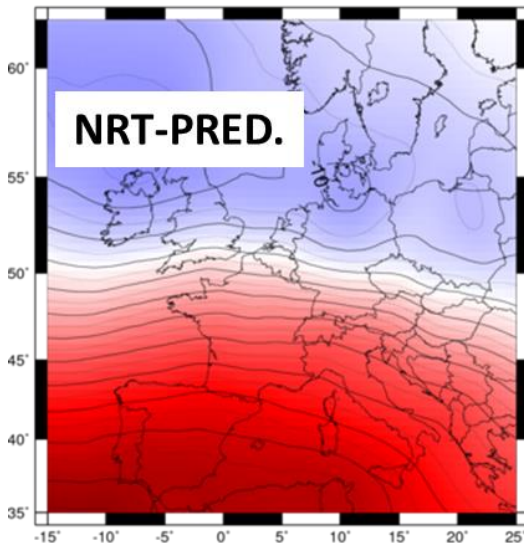


Median VTEC last 15 days 18:15-18:30 UTC

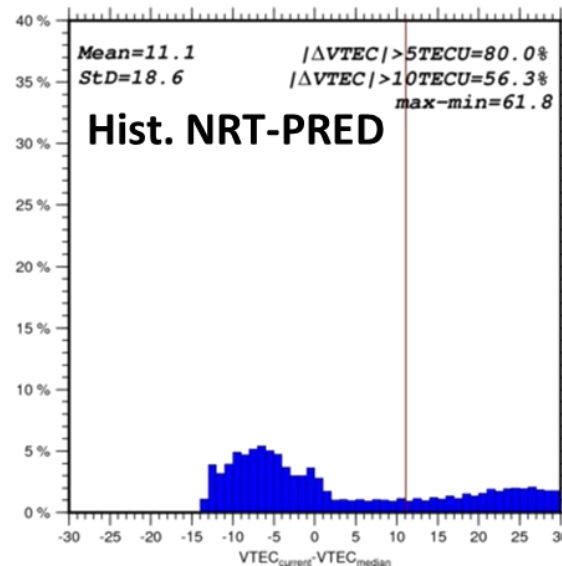


Normal ionospheric TEC behaviour : median of the VTEC for the 15 previous days

Difference $VTEC_{current} - VTEC_{median}$

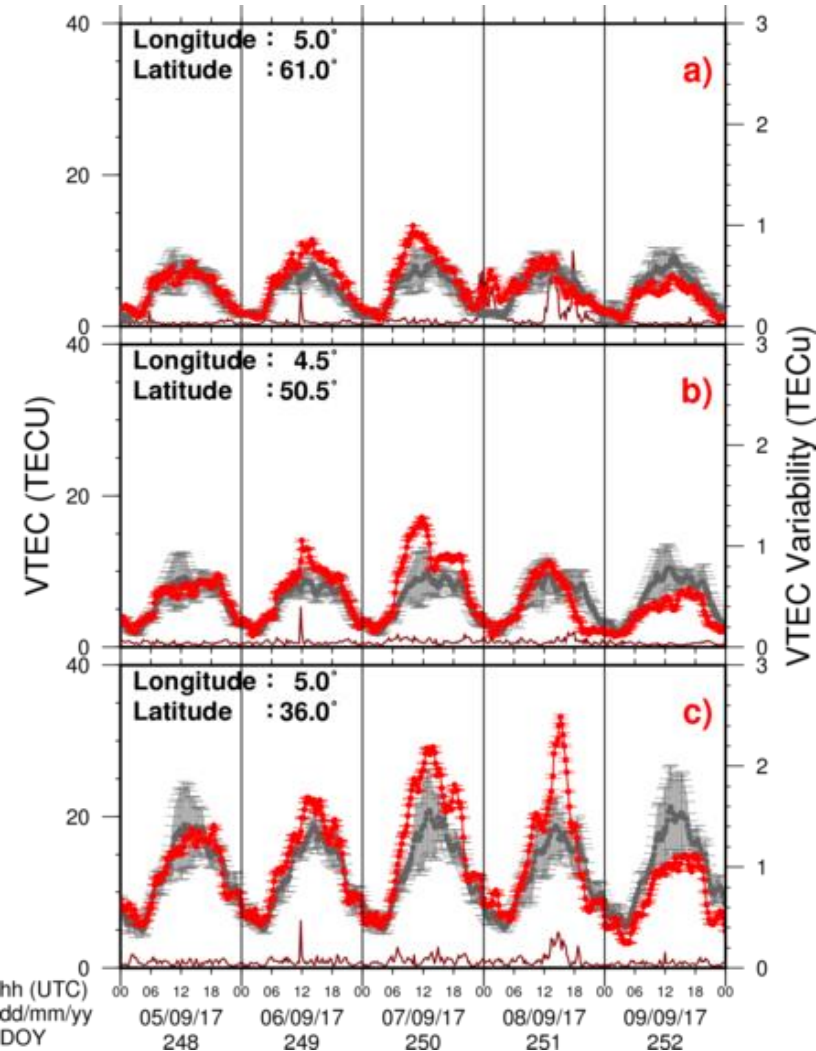


Histogram of the differences



*Saint Patrick storm
March 17th 2015 event*

SUMMARY OF THE EVENT: A solar flare occurred the 6/09/2017 generating a sudden small increase of TEC at noon with higher variability of TEC. The next days, the arrival of the CME generated disturbances in the North during night-time of the 7/09/2017 and at the end of the day 07/09/2017. An increase of TEC was also observed in the South the 07 and 08/09/2017. A depletion of TEC followed the following day 09/09/2017



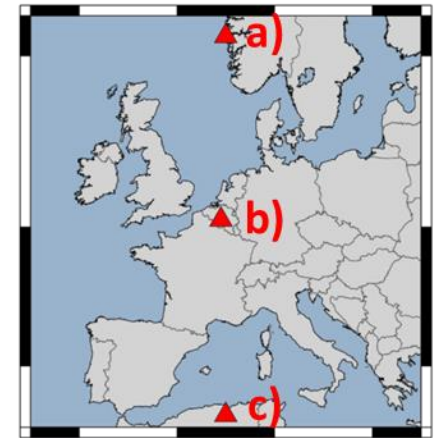
TEC extracted from the NRT maps

TEC current

15-days median TEC

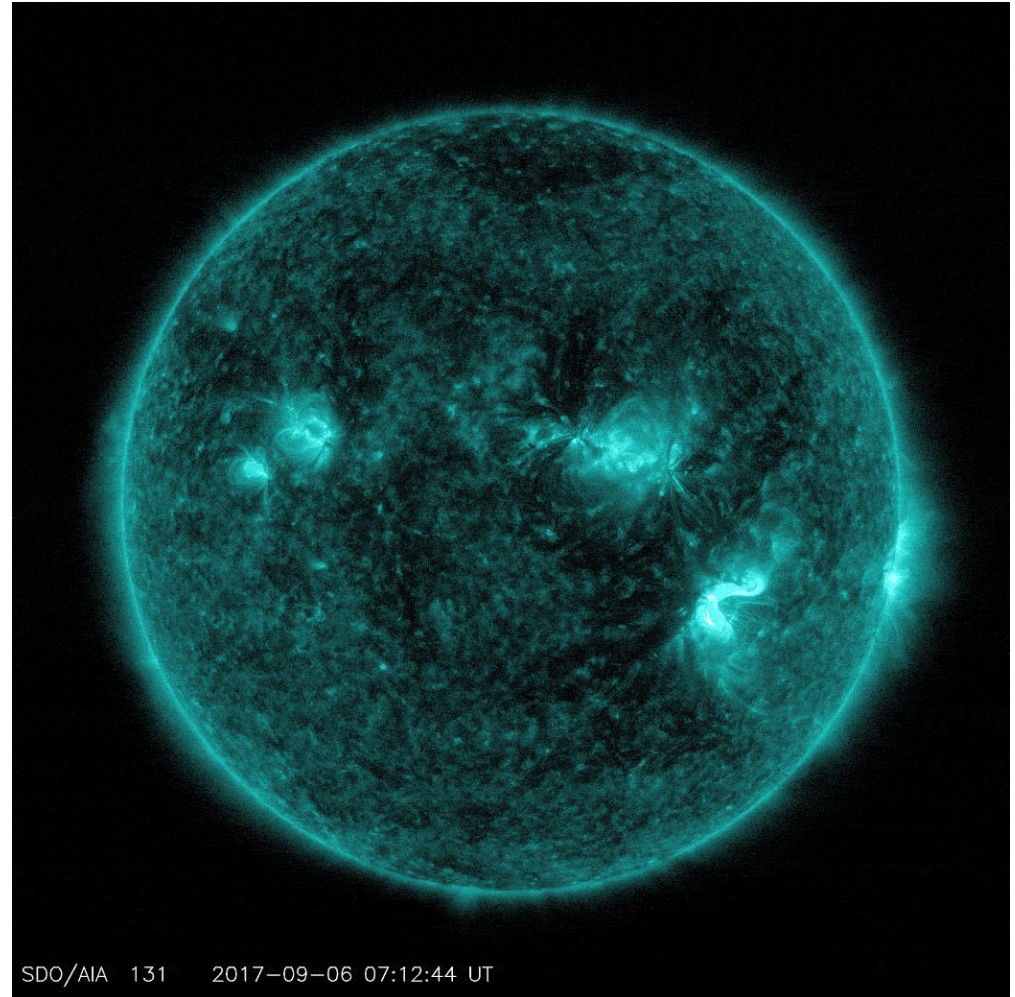
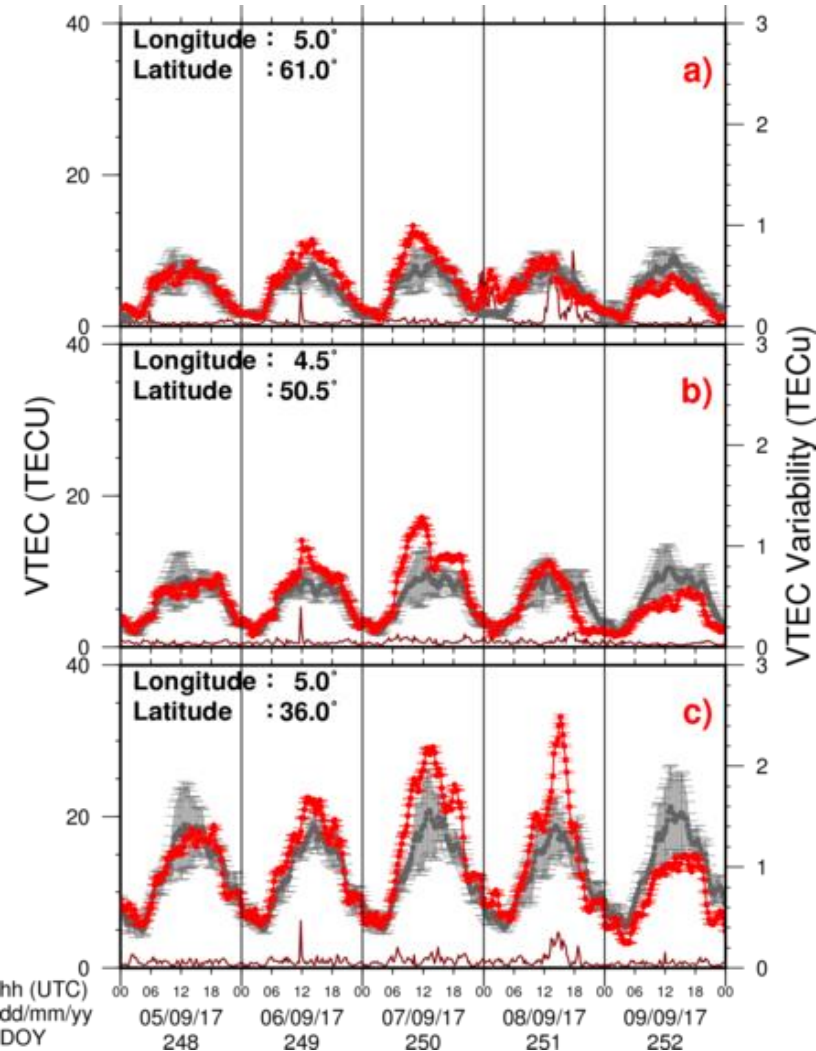
TEC Variability

+ Links to interactive and statistical maps, and origin of the event sidc.oma.be

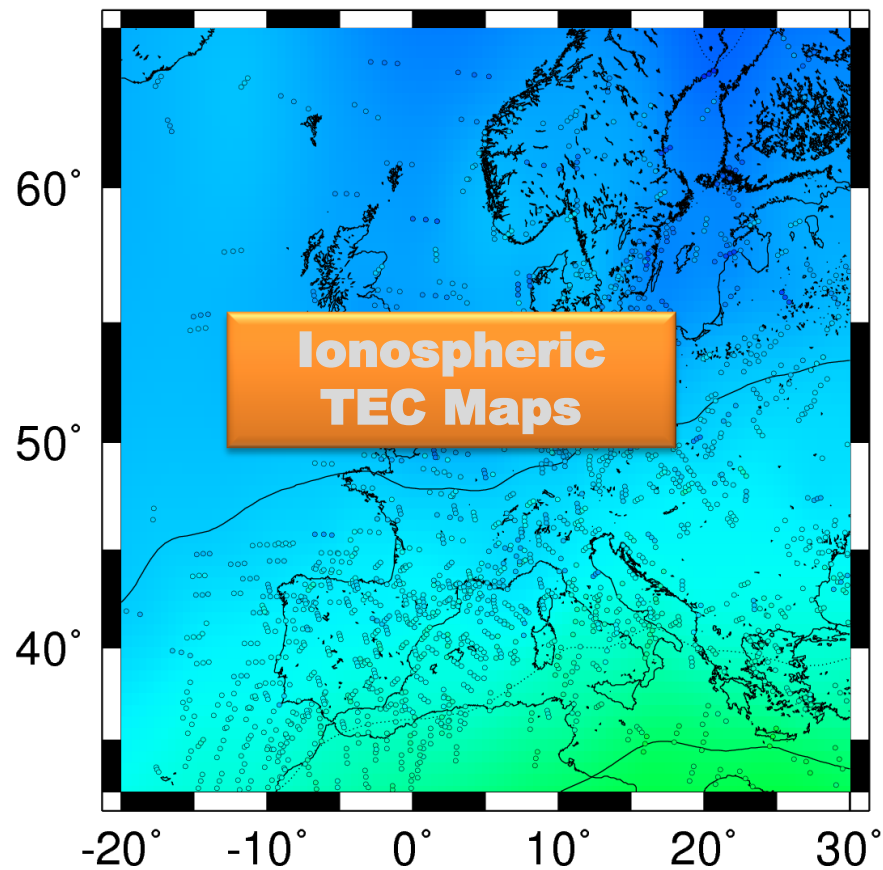


2012–2017 (43 events)

SUMMARY OF THE EVENT: A solar flare occurred the 6/09/2017 generating a sudden small increase of TEC at noon with higher variability of TEC. The next days, the arrival of the CME generated disturbances in the North during night-time of the 7/09/2017 and at the end of the day 07/09/2017. An increase of TEC was also observed in the South the 07 and 08/09/2017. A depletion of TEC followed the following day 09/09/2017



Research and Scientific Applications



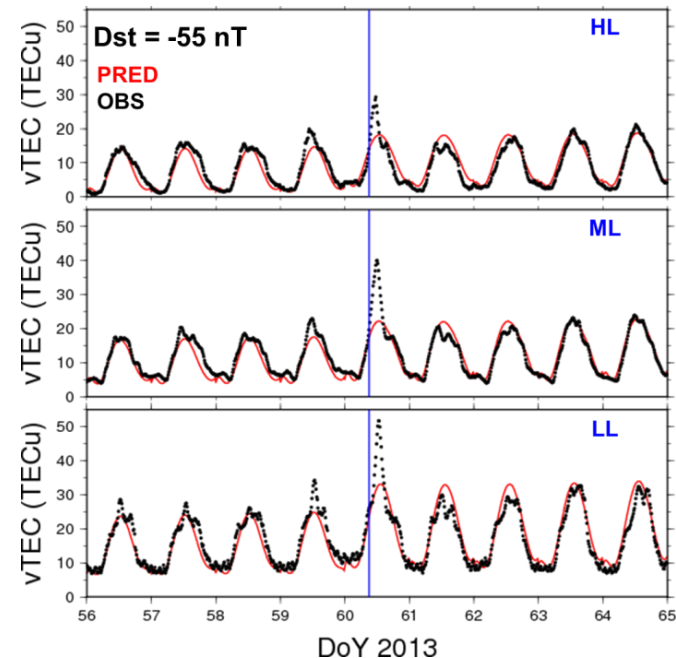
Climatology of the Ionosphere

Empirical Model, least-square adjustment with :

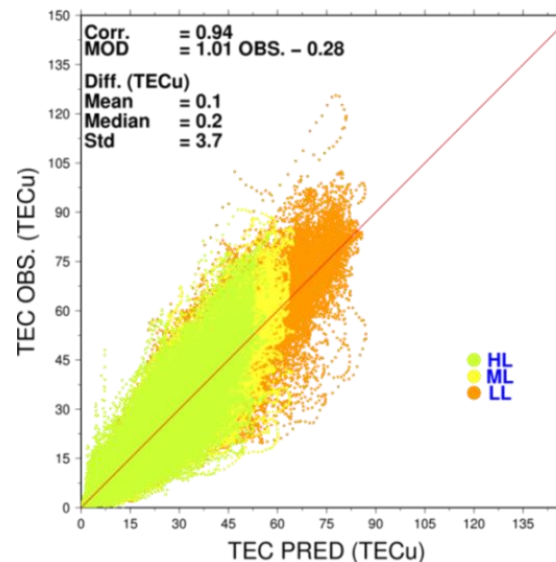
- 8th order polynomial function with monthly coefficients between the TEC and F10.7P
- Discretization with respect to the solar activity phases

Bergeot et al. 2015, EGU

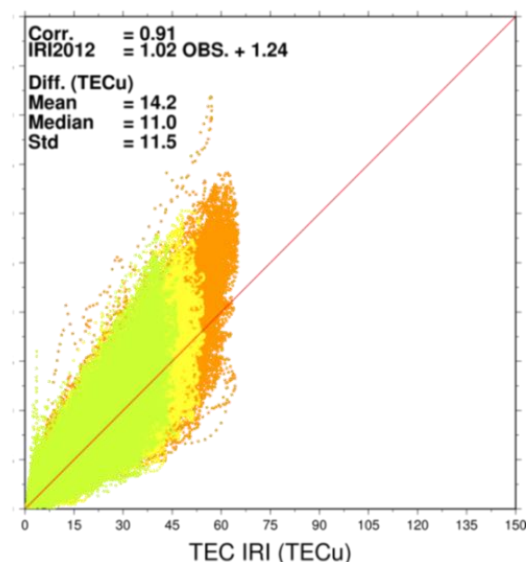
Ionospheric TEC Maps



GNSS-based model scattering plot



IRI model scattering plot



Research and Scientific Applications

Climatology of the Ionosphere



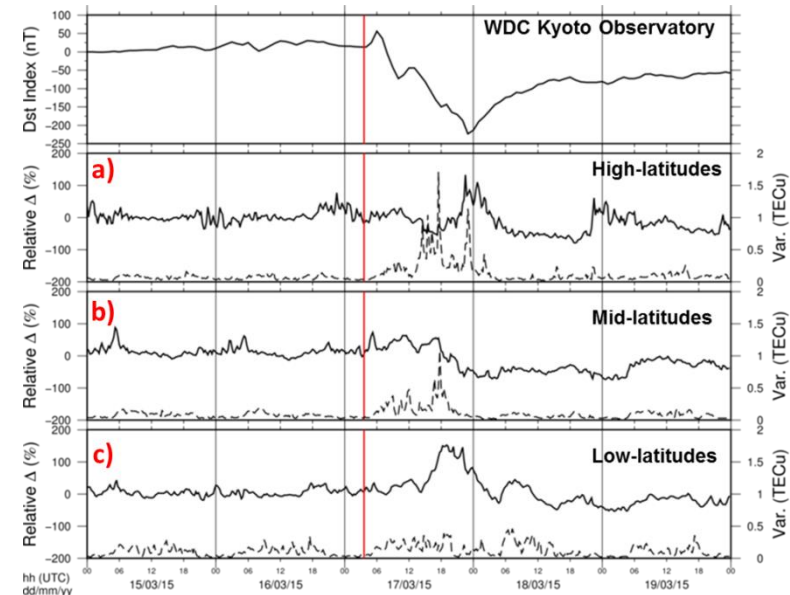
IAG – Real-Time Ionospheric Monitoring Working Group

Comparison of current Ionospheric Models:

March 17, 2015 - St Patrick Storm

Ionospheric TEC Maps

Garcia-Rigo et al. 2017, EGU



Climatology of the Ionosphere



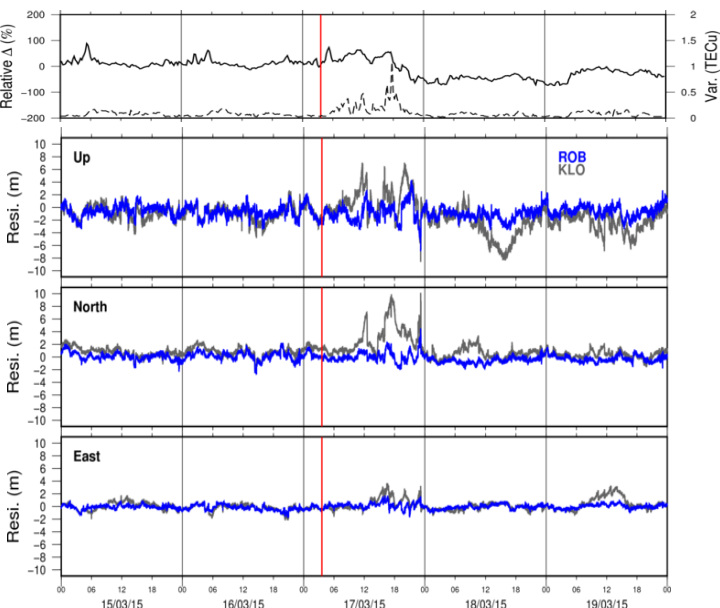
IAG – Real-Time Ionospheric Monitoring Working Group

Improving GNSS single frequency positioning

Position of the GNSS station at Brussels during 2015 March Storm (*W. Huang and P. Defraigne*)

Bergeot et al. 2015, URSI

Ionospheric TEC Maps



Correction using Klobuchar ionospheric model

| | |
|-------|------------------|
| East | 10 ± 80 cm |
| North | 100 ± 140 cm |
| Up | 120 ± 210 cm |

Correction using ROB-TEC products

| | |
|-------|-----------------|
| East | 6 ± 40 cm |
| North | 9 ± 66 cm |
| Up | 76 ± 150 cm |

Climatology of the Ionosphere

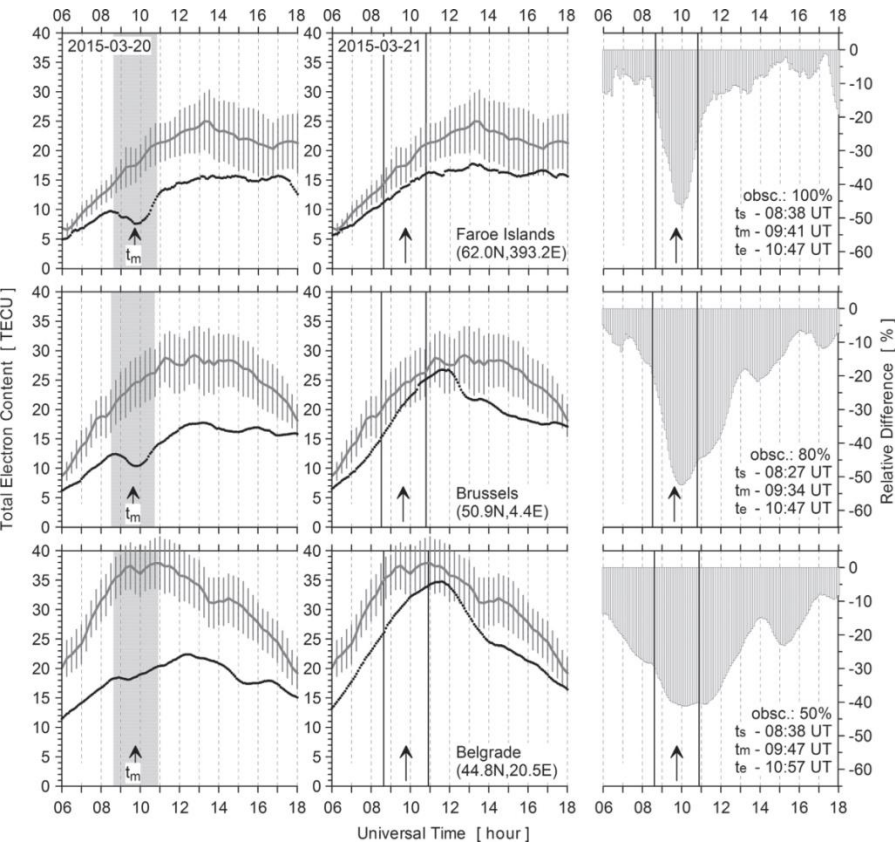
Improving GNSS single frequency positioning



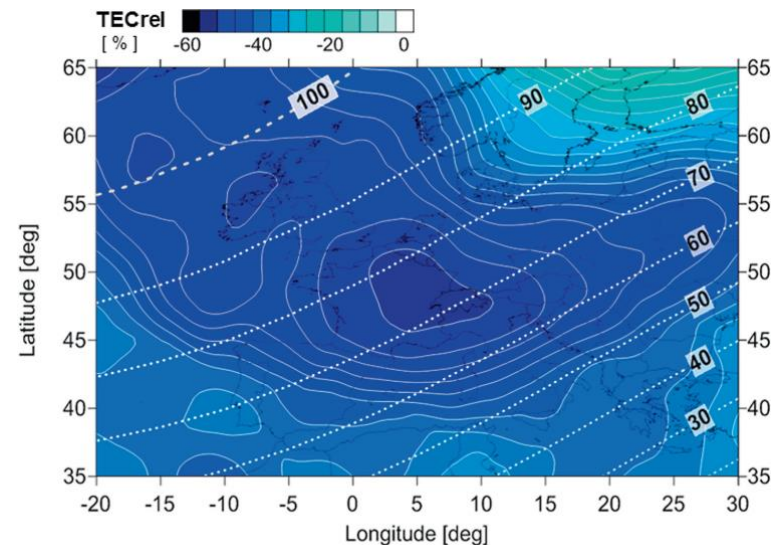
IAG – Real-Time Ionospheric Monitoring Working Group

Solar Eclipse 20th Mar. 2015 RT and post

Ionospheric TEC Maps



Stankov et al. 2017, SWSC Journ.



Research and Scientific Applications

Climatology of the Ionosphere

Improving GNSS single frequency positioning

3D Ionosphere Nowcasting and Forecasting for ESA Space Situational Awareness



National observatory of Athens



ESA

Belehaki et al. 2015, SWSC Journ.

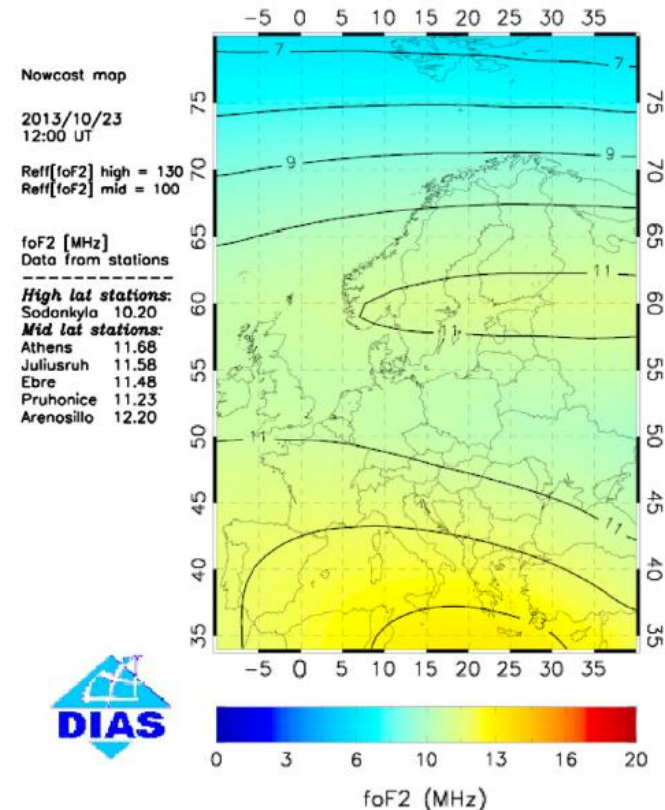
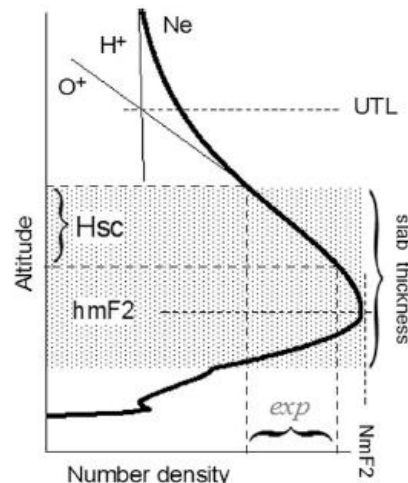


IAG – Real-Time Ionospheric Monitoring Working Group

Solar Eclipse 20th Mar. 2015 RT and post

Ionospheric TEC Maps

with Digisonde Network Data



Research and Scientific Applications

**Climatology of
the Ionosphere**

**Improving GNSS single
frequency positioning**



**IAG – Real-Time
Ionospheric Monitoring
Working Group**

**Solar Eclipse 20th Mar. 2015
RT and post**

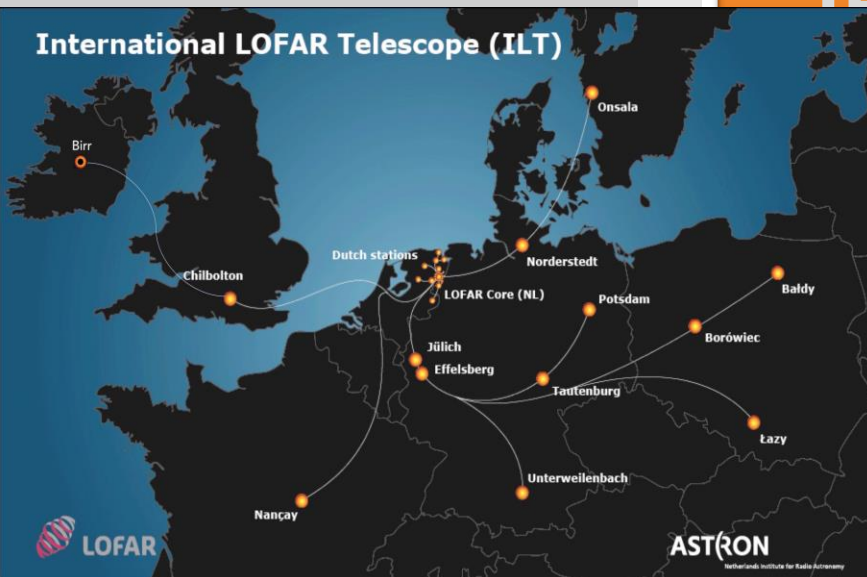
**Ionospheric
TEC Maps**

**Calibration of LOFAR
radio telescope**

***Sotomayor-Beltran et al. 2013 Astronomy
& Astrophysics***

Use of ROB-TEC maps to remove the time-
variable ionospheric Faraday rotation
contribution

International LOFAR Telescope (ILT)



**Climatology of
the Ionosphere**

**Improving GNSS single
frequency positioning**

**3D Ionosphere
Nowcasting and
Forecasting for
ESA Space Situational
Awareness**

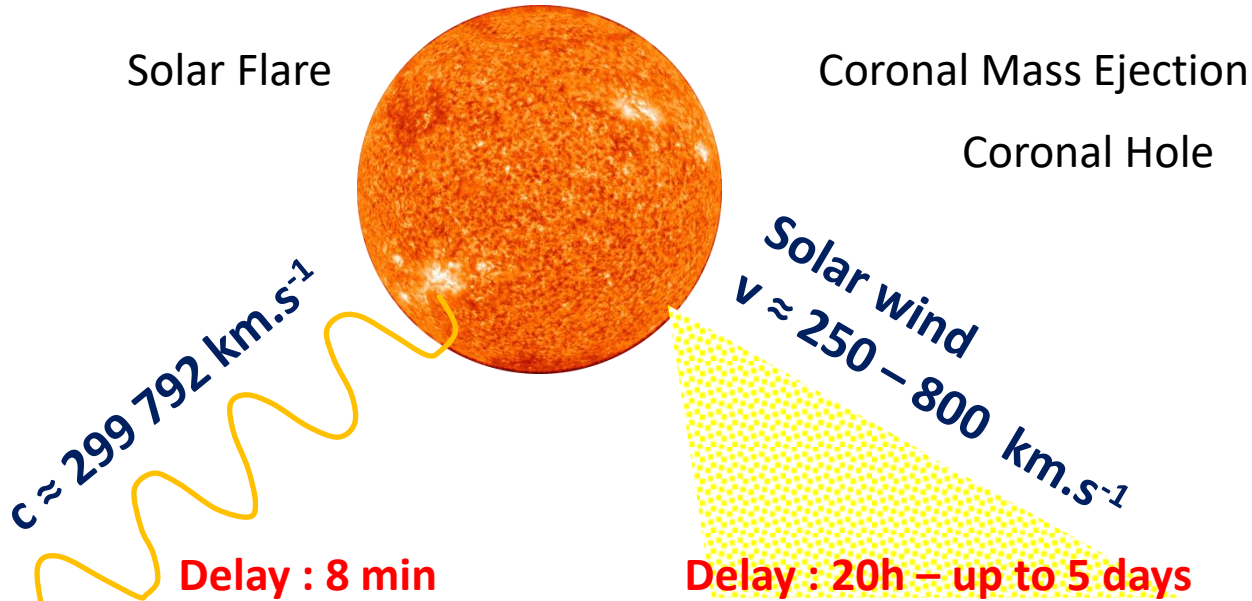


**IAG – Real-Time
Ionospheric Monitoring
Working Group**

**Solar Eclipse 20th Mar. 2015
RT and post**

**Ionospheric
TEC Maps**

**Calibration of LOFAR
radio telescope**



| <u>Electromagnetic Radiations</u> | |
|-----------------------------------|-------------|
| EUV X-rays | Radio Waves |

Photo-ionisation

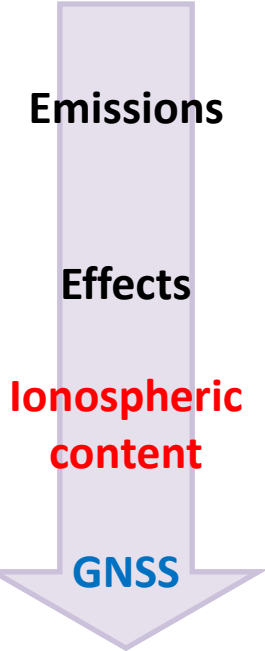
Solar Radio Burst

Ionospheric Ne Δ - Scintillations

Ionospheric content

Radio signal delay - \searrow Signal reception

Signal reception fade (\nearrow noise)



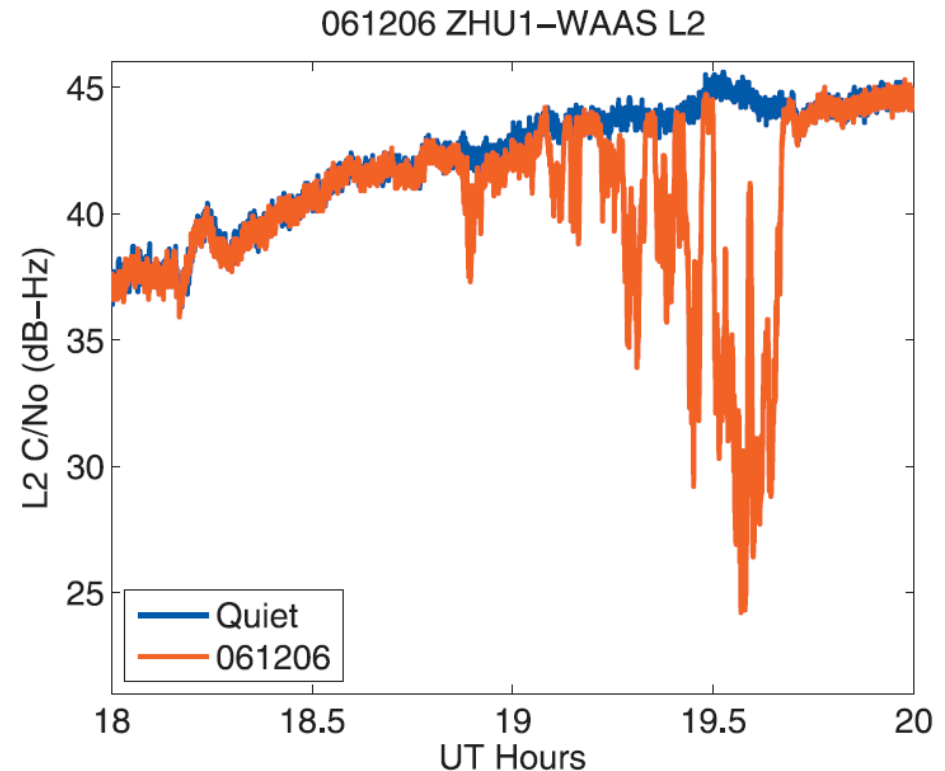
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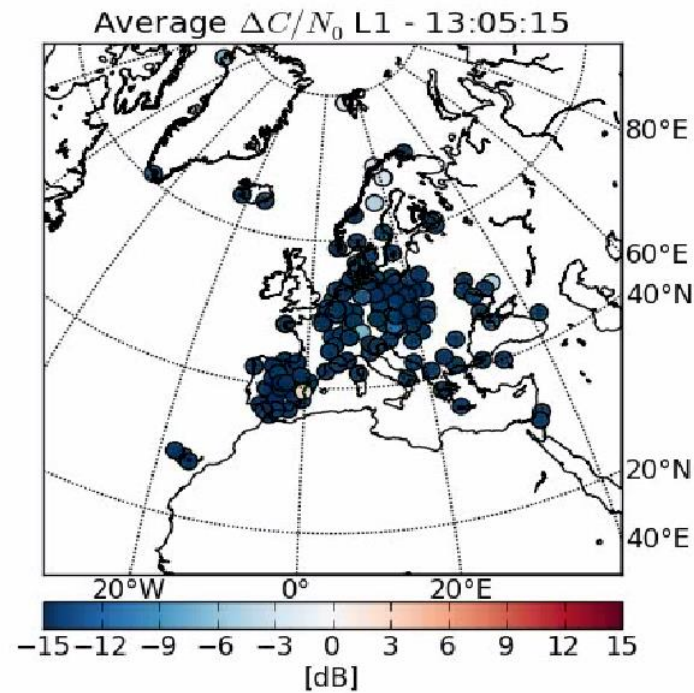
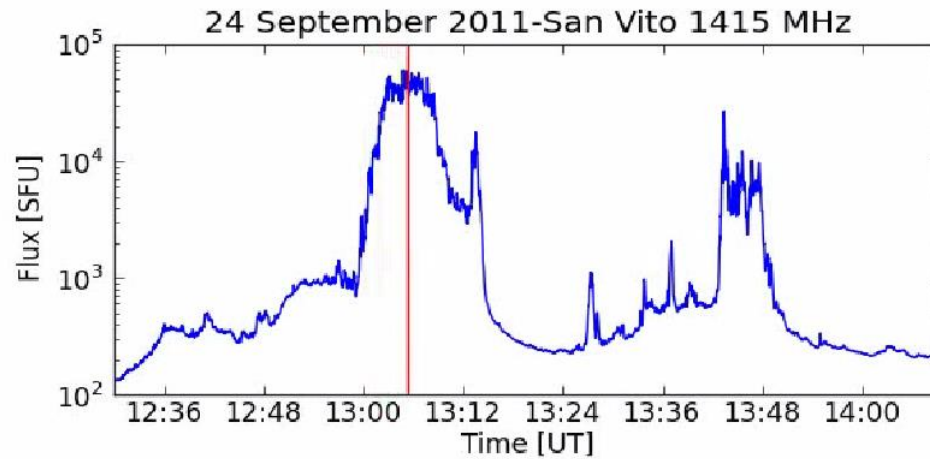
Geomagnetic Storm
Ionisation & Recombination - Aurora

Ionospheric Ne disturbances - Scintillations

Radio signal delay - \searrow Signal reception

- The Sun emits in radio over a wide frequency range (from few kHz to GHz)
- Solar Radio Bursts (**SRB**) are intense radio emissions (durations from 10s to few hours)
- SRBs increase the noise level of GNSS ground stations
- **Carrier-to-Noise density (C/N_0)** [35; 55] dB-Hz





RT monitoring of the abnormal fade of GNSS signal reception due to SRB at the 2 GNSS frequency bands L1 and L2

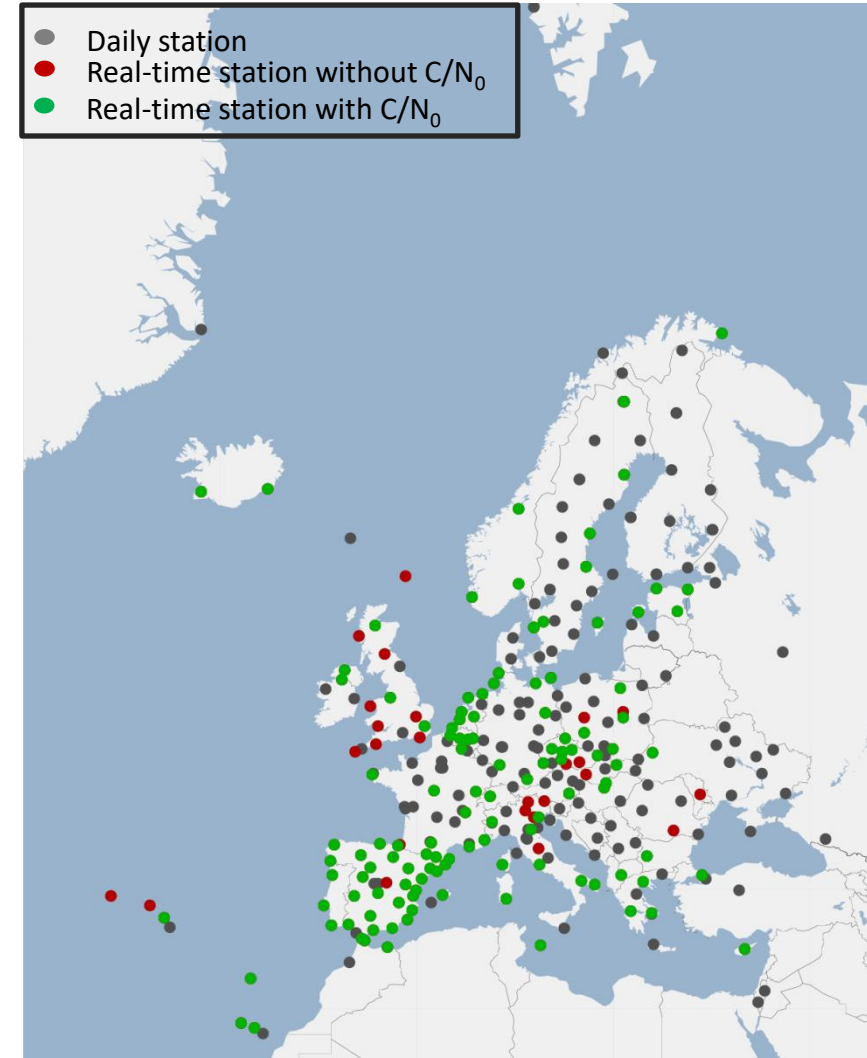
Chevalier et al., URSI GASS 2017

C/N_0 (dB-Hz) extracted from RINEX files (S1-S2)

- But no standardized unit

⇒ **Please provide C/N_0 instead of Signal to Noise Ratio (SNR) (manufacturer/receiver dependent)**

| Level | GNSS $\Delta C/N_0$ Fade | Effect |
|----------|-----------------------------|--|
| Quiet | >-1dB-Hz | none |
| Moderate | -1 dB-Hz | SRB detected but should not impact GNSS applications |
| Strong | -3 dB-Hz | Potential impact on GNSS applications |
| Severe | -10 dB-Hz | Potential failure of the GNSS receivers |



Last Ionospheric Events

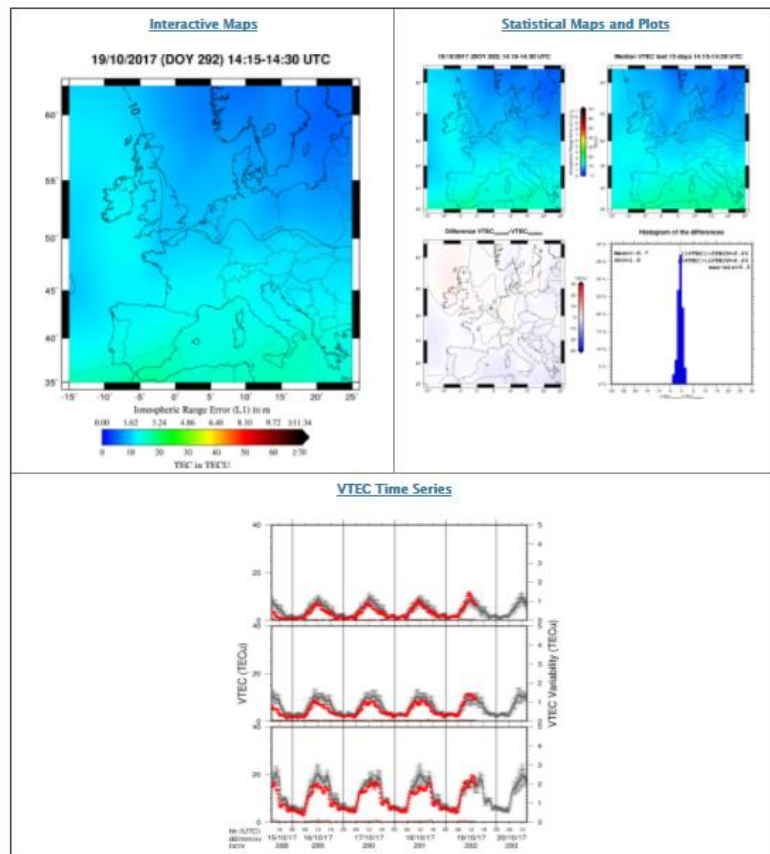
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- ... [more events here](#)

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Vertical Total Electron Content (VTEC) estimated in Near Real-Time (NRT) every 15 minutes from EUREF Permanent Network (EPN) GPS data.

[More...](#)

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www.gnss.be

1) Ionospheric Total Electron Content (TEC)

- Interactive TEC maps
- Statistical TEC maps
- TEC Time Series at 3 locations (North-Brussels-South)
- TEC Data (IONEX)
<ftp://gnss.oma.be/gnss/products/IONEX/>
- Event Description

2) Solar Radio Burst (SRB)

- Warning System
- Event Description

Solar Radio Burst Warnings for GNSS Applications in Europe

Solar Radio Bursts (SRB) emitted at the GNSS frequencies can affect the GNSS signal reception. To detect such event, a [near-real time SRB warning system](#) with a 4-level index was set in Europe using the real-time EUREF Permanent Network.

www.gnss.be

SRB WARNING SYSTEM FOR GNSS APPLICATIONS IN EUROPE

Contact: iono@oma.be

To receive real-time alert emails, please contact us to be added to the mailing list.

| Last update : 2015-11-04 14:29:30 | L1 | L2 |
|-----------------------------------|----|----|
| Last 15min | | |
| Last 24h | | |
| Last week | | |

Events of the last 30 days:

| Frequency | Date of the maximum fade | Maximum fade (in dB/Hz) | Beginning of the event (fade < -1 dB/Hz) | End of the event (fade > -1 dB/Hz) |
|-----------|--------------------------|-------------------------|--|------------------------------------|
| L2 | 2015-11-04 14:29:00 | -5.75±2.26 | 2015-11-04 14:22:00 | On-Going |

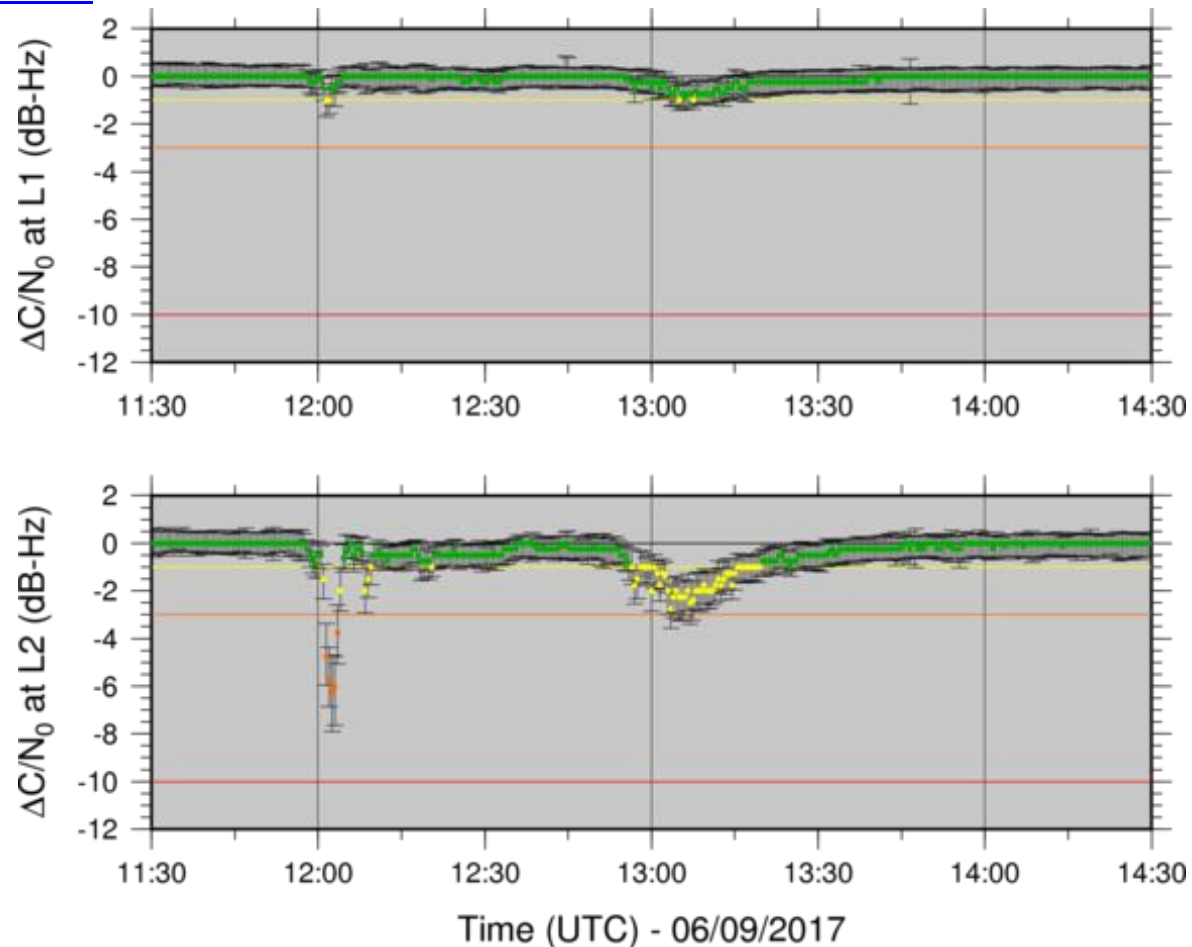
To detect Solar Radio Burst (SRB) affecting the GNSS signal reception in Europe, the carrier to noise density (C/N0) of the real-time EPN GPS network are monitored in near-real time (updated every 15 minutes). The intensity of the SRB impacts on GNSS applications are indexed at 4 levels:

- **quiet**
- **moderate** : SRB detected but should not impact GNSS applications,
- **strong** : potential impact on GNSS applications,
- **severe** : potential failure of the GNSS receivers.

Previous SRB Events at the GNSS frequencies

- 2015-11-04 : Space weather event due to Solar Radio Burst ([more here](#))

SUMMARY OF THE EVENT: The solar radio bursts of the 06/09/2017 impacted the GPS signal reception at both frequencies L1 and L2. On L1, two fades above 1dB-Hz were detected at 12h01 and 12h05. On L2, a first fade above 3dB-Hz which could potentially affect the GNSS application, occurred for 3 min with a maximum of -6.25 ± 1.6 dB-Hz at 12h02. It was followed by a second lower fade above 1dB-Hz at 13h03. For additional information about the burst on a larger frequency spectrum see at [SIDC Humain Radioastronomy Station](#).



■ IONOSPHERE

- TEC maps (+TEC variability) over Europe since 2012 in near-real time (0.5°x0.5° grids, 15 min.)
- Visualisation of the ionospheric activity : www.gnss.be
- Data: <ftp://gnss.oma.be>
- Scientific applications : Climatology of the Ionosphere, GNSS single frequency positioning(...)

■ SOLAR RADIO BURST

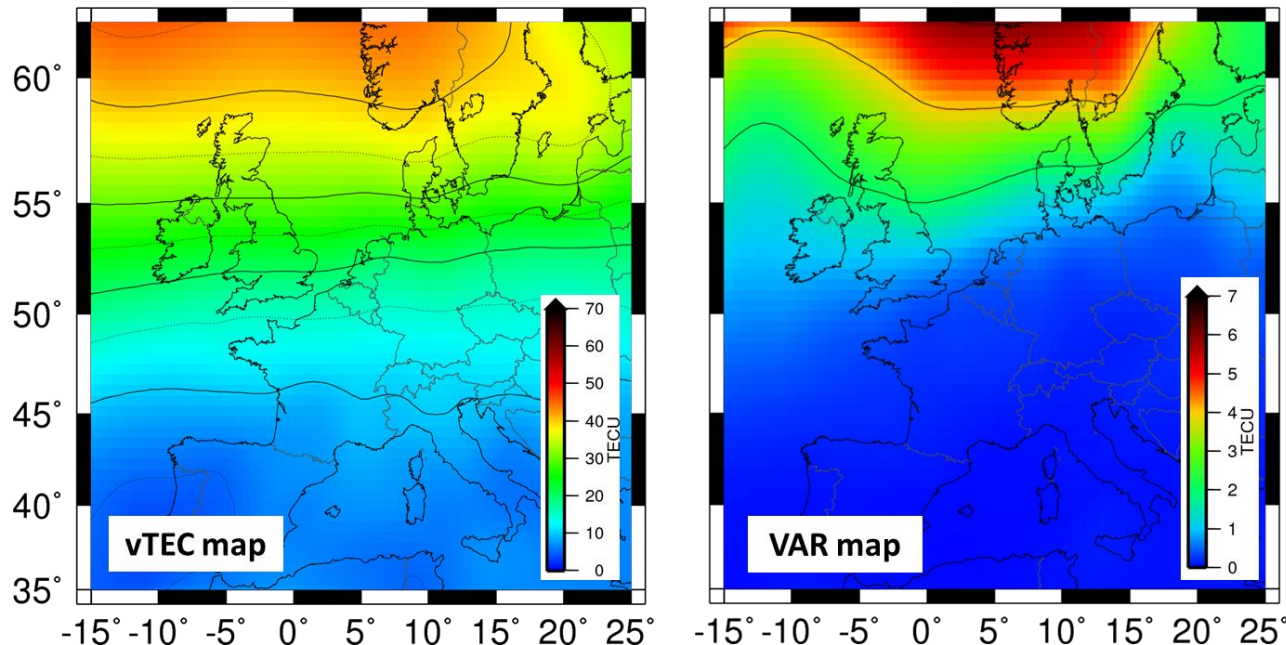
- GNSS signal reception is monitored in real-time
- SRB Warning System
- Register at the email alert : iono@oma.be

Thank you

Back-up slides

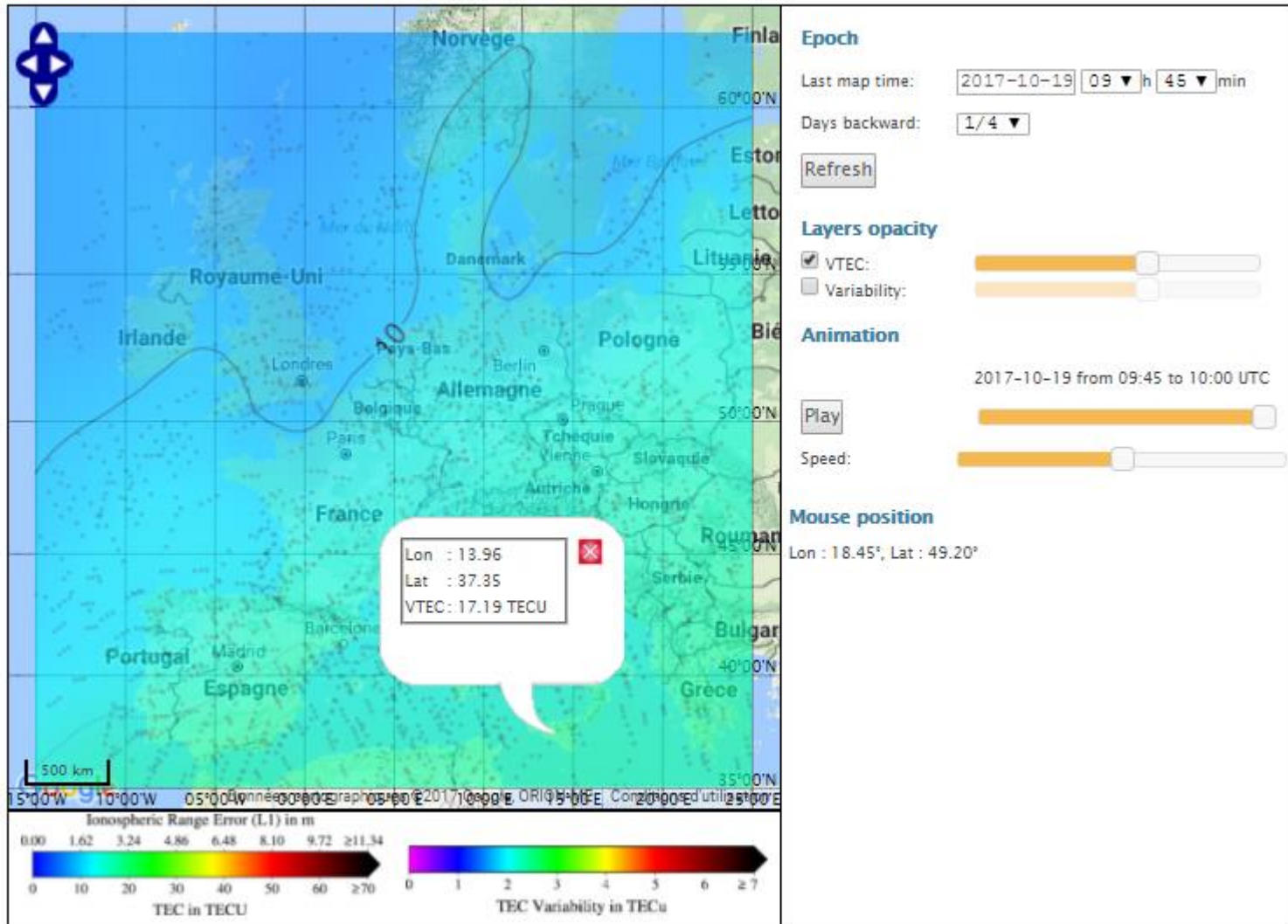
- Post processing – all available GPS+GLONASS data
- Time independent, tested and validated during quiet time and, minor and major events

HALLOWEEN STORM 2003 (22:15-22:30 UTC)



Contact: iono@gnss.be

2017-10-19 (day 292) from 09:45 to 10:00 UTC



Click on the map to get the Vertical Total Electron Content (VTEC) value at a particular point.

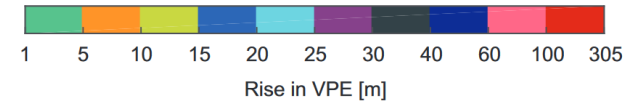
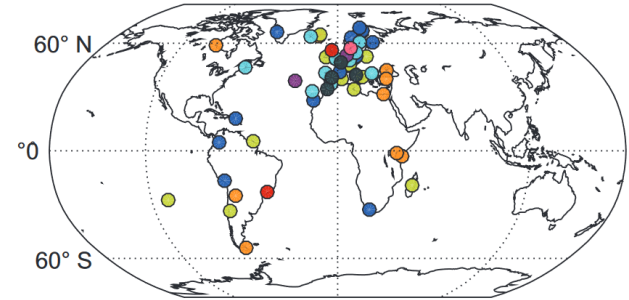
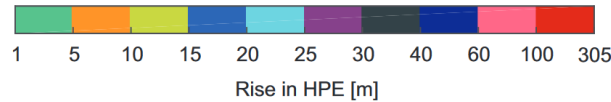
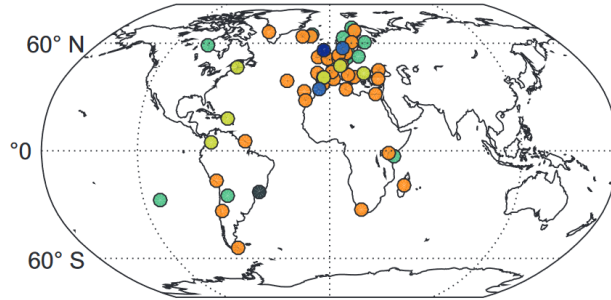
Rise in error = Positioning error during SRB – Positioning error on a quiet day

120° W 0° 120° E

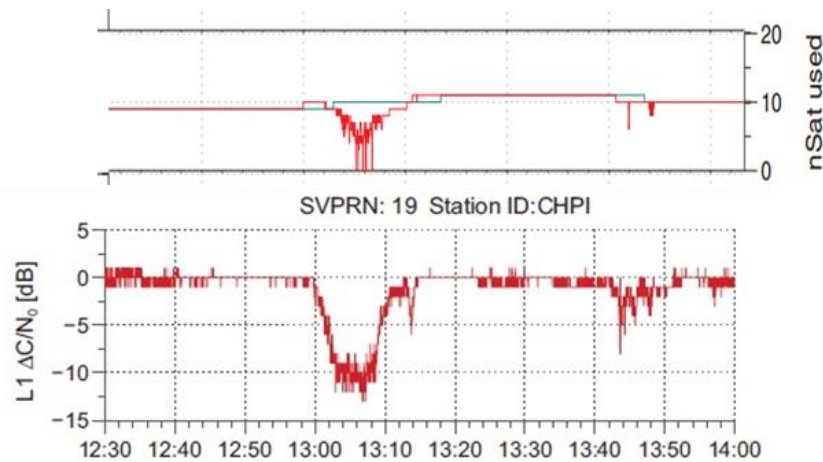
120° W 0° 120° E

**SRB of the
24/09/2011**

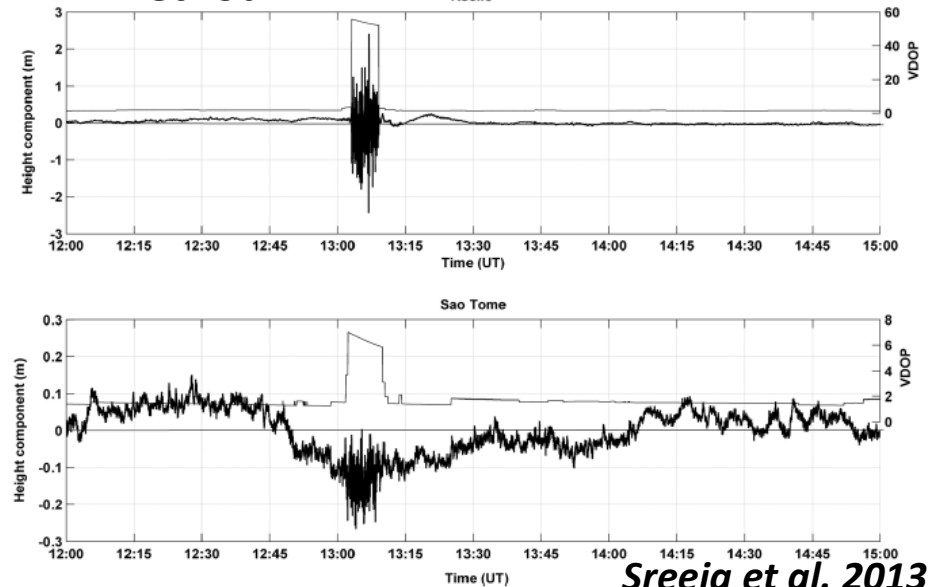
**Muhammad et
al. 2015**



Sao Paulo, Brazil



PPP method



Sreeja et al. 2013